Progress (...) on forward modelling L1 and L2 bending angles

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Motivation

• How to do ionospheric correction when L2 signal drops out?

• Different (better?) tropospheric retrievals if L1 and L2 both used?

• Some information on ionospheric parameters?

• Residual ionospheric error?



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Model ionosphere: electron density



Single Chapman Layer (Chapman 1931) $n_e(r) = n_e^{max} exp[\frac{1}{2}(1 - u - e^{-u})],$ where $u = (r - r_0) / H$. 3 parameters: $n_e^{max} = TEC / \sqrt{(2\pi e)} H$ $r_0 = peak height$

H = ionospheric scale height

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Model ionosphere: bending angle

 $\alpha_{\text{Li}}(a) = -2a \int_a^{\infty} d\log n/dx \, dx \, / \, \sqrt{(x^2 - a^2)}, \quad x = nr$

 $\approx (k_4 / f_{Li}^2) 2a \int_a^{\infty} dn_e / dx dx / \sqrt{(x^2 - a^2)}, \ k_4 = 40.3 \text{ m}^3 \text{s}^{-2}$

≈ (k₄ / f_{Li²}) 2a [2r₀ / (r₀+a)^{3/2}]
$$\int_a^{\infty} dn_e / dr dr / \sqrt{(r - a)}$$

= $(k_4 / f_{Li}^2) n_e^{max} [4er_0^2 a^2 / H(r_0 + a)^3]^{1/2} \cdot Z(\ell)$

where

 $Z(\ell) = \int_{-\ell}^{\infty} (e^{-u} - 1) \exp[\frac{1}{2}(1 - u - e^{-u})] / \sqrt{(u + \ell)} du$

is just a dimensionless, O(1) function of

 $\ell = (\mathbf{r}_0 - \mathbf{a}) / \mathbf{H}$





Model ionosphere: bending angle $Z(\ell)$

Schreiner et al, 1999, Chapman layer

Melbourne, 2004, Gaussian n_e

Hajj & Romans, 1998, GPS/MET



Direct modelling of L1 and L2 in ROPP

- The Radio Occultation Processing Package:
 - A collection of Fortran 95 code, build and test scripts, data files and documentation designed to aid users wishing to process, quality-control and assimilate radio occultation data into their NWP models.
 - Provided by ROM SAF (EUMETSAT).
 - The following features will be available in ROPP8.0 (Dec 2014).



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ROM SAF Workshop, ECMWF, 16-18 June 2014

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Direct modelling of L1 and L2 in ropp_1dvar

- Minimise $2J(\mathbf{x}) = (\mathbf{x} \mathbf{b})^{\top} \mathcal{B}^{-1} (\mathbf{x} \mathbf{b}) + (\mathcal{H}(\mathbf{x}) \mathbf{o})^{\top} \mathcal{R}^{-1} (\mathcal{H}(\mathbf{x}) \mathbf{o})$
- $\mathbf{x} = \{T, q, p^*, n_e^{max}, r_0, H\};$ $\mathbf{o} = \{\alpha_{L1}, \alpha_{L2}\}$
- · R
 - $\sigma(\alpha_{L1}) = \max(\alpha_n, 10 \ \mu rad); \ \sigma(\alpha_{L2}) = \max(\alpha_n, 30 \ \mu rad)$
 - Assume α_{L1} and α_{L2} errors to be uncorrelated
 - Needs some experimentation
- · B
 - $\sigma(n_e^{max}) = 2e11 \text{ m}^{-3}; \sigma(r_0) = 150 \text{ km}; \sigma(H) = 25 \text{ km}$
 - Assume {n_e^{max}, r₀, H} errors to be uncorrelated from each other and from those of {T, q, p*}
 - Needs some experimentation



Example 1: ropp_1dvar retrieval based on neutral bending angle



Example 1: ropp_1dvar retrieval based on L1 & L2 bending angles



Example 2: ropp_1dvar retrieval based on neutral bending angle



Example 2: ropp_1dvar retrieval based on L1 & L2 bending angles



Example 3: ropp_1dvar retrieval based on neutral bending angle



Example 3: ropp_1dvar retrieval based on L1 & L2 bending angles



Example 4: ropp_1dvar retrieval based on neutral bending angle



Example 4: ropp_1dvar retrieval based on L1 & L2 bending angles



Direct modelling of L1 and L2 in ROPP: α_n stats



Direct modelling of L1 and L2 in ROPP: α_{Li} stats



"Retrieved ionospheric parameters"

Fit "normal" L1-L2 to $Z(\ell)$ for every GRAS occultation for 5 days in June 2012

TEC (TECU)

GRAS 11/06/2012 - 15/06/2012



Direct modelling of L1 and L2 in ROPP: summary

- Can forward model α_{L1} and α_{L2} by assuming simple Chapman layer ionosphere
- Allows extrapolation of α_{L1} and α_{L2} based on whole profiles of data
- Retrievals using α_{L1} and α_{L2} are possible in principle BUT
 - Bigger impacts on q and tropospheric T?
 - Not always capturing curvature of L1 and L2 at top?
 - Require $n_e^{max} < 0$ to get decent fit in some cases!
- Backgrounds/errors/covariances of n_e^{max}, r₀ and H need some experimentation.
- Errors/covariances of α_{L1} and α_{L2} need some experimentation.



Residual ionospheric correction

- Working on a model for the residual error that remains after standard ionospheric correction (Vorob'ev and Krasil'nikova,1994)
- $\alpha(a) = -2a \int_a^{\infty} d\log n/dx \, dx \, /\sqrt{(x^2 a^2)} \approx$

 $-2a \int_{a}^{\infty} N' dr / \sqrt{(r^2 - a^2)}$ Basic bending angle, $\propto TEC/f^2$

+ $2a \int_{a}^{\infty} NN' dr (2r^2 - a^2)/(r^2 - a^2)^{3/2}$ 1st order correction, $\propto TEC^2/f^4$

- Implies that residual after standard ionospheric correction

 ∝ (α_{L1} − α_{L2})² X some *slowly* varying function of height,
 which we can estimate/calculate analytically for a Chapman layer.
- We are working with Julia Danzer (GFZ/ROM SAF VS) to test this.
- Could have implications for RO-based climatological mean temperatures.





Conclusions

 Calculated bending angle produced by a spherically symmetric Chapman layer ionosphere.

Incorporated in ROPP forward model and 1dvar system.

• Expressed residual bending angle, after standard ionospheric correction, in terms of $(\alpha_{L1} - \alpha_{L2})^2$.



Background error covariance matrix



Bending angle error covariance matrix



Fit "normal" L1-L2 to $Z(\ell)$ for every GRAS occultation for 5 days in June 2012

r₀ - R₌ (km)

GRAS 11/06/2012 - 15/06/2012



Fit "normal" L1-L2 to $Z(\ell)$ for every GRAS occultation for 5 days in June 2012

