IASI PC compression – Searching for signal in the residuals

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Lossless compression = Noise preserving compression

SIGNAL

spectral redundancy \rightarrow exploited in the current PC compression spatial redundancy \rightarrow could be exploited to compress further

RANDOM NOISE

no spectral redundancy and no spatial redundancy

➔ no compression possible, except entropy encoding (like Huffman coding)

Noise constitutes the bulk of the "information" in the spectra

insisting on preserving instrument noise imposes strict limits to the obtainable compression factor - the limit depends on the chosen quantisation step



PCR quantisation compared to L1C quantisation

Essence of EUMETSATs IASI PC compression

- A. Band separation
 - compression of Band 1 and 2 still possible when Band 3 is bad
 - inter band correlation not exploited ightarrow need 40 PC scores extra
- B. Diagonal noise normalisation matrix (N)
 uniformises but doesn't decorrelate the noise
- C. Eigenvectors built from training set of real measurements (Y)- outliers added to capture rare situations like fires and volcanic eruptions
- D. Number of retained eigenvectors (90, 120, 90) based on spatial correlation of PC scores
 - $\rightarrow E$ a truncated set of eigenvectors of COV($N^{-1}Y$)

 $p = E^T N^{-1}(y - \overline{y})$ PC scores $\widetilde{y} = NEp + \overline{y}$ Reconstructed radiances

 $\widetilde{y} = Ay + (I - A)\overline{y}$ where $A = NEE^T N^{-1}$

The transformation to reconstructed radiances is a projection!

The most important slide! (or how to split a spectrum into four parts)



Atmospheric signal retained in reconstructed radiance Instrument noise retained in reconstructed radiance Atmospheric signal in residual (RECONSTRUCTION ERROR ⁽²⁾) Instrument noise in residual the <u>covariance of the residuals</u> is the sum of the <u>noise in the residuals</u> and the <u>covariance of the reconstruction error</u>

Raw and reconstructed noise covariance matrices:



Estimation of full noise covariance matrix





Estimation of full noise covariance matrix

Additional noise correlation caused by spectral calibration difficulties in the band overlap regions.



Monitoring of noise evolution

Decontamination, September 2010.

Noise reduction of 0 -20% in each channel.

Agrees well with black body based estimates of the noise reduction





Outliers

spectra which don't reconstruct well

RMS of noise normalised residual (reconstruction score)

- one for each IASI band
- disseminated with the IASI PC score product



Kasatochi eruption (Band 2 Residual RMS)



Photonic noise - increases with the signal

The noise and therefore the expected value of the residual RMS depends on the radiance sum and the detector \rightarrow

This must be taken into account to get a sensitive detections of outliers.



Scatter plots of residual RMS vs. radiance sum (Band 2)

Residual_RMS - slope*Radiance_Sum > Threshold(detector)

outlier spectra gathered in an auxiliary product (IASI_IPO)

Russian fires, August 3. to 12. 2010

800

3.0 -

÷ 2.0 -2.0 -1.0 -

0.0

640

a total of 73 outliers collected in this area



900

10

Ammonia (NH)



Residual statistics (std and mean) for band 1 outliers over Russia, August 3-12 2010 (73 spectra.)

What causes the remaining outliers?

Undetected "spikes":

High-frequency disturbance of the interferogram, most often observed in the South Atlantic Anomaly. (Band 2 outliers)

Back to normal operation after external calibration mode:

No history available for deriving filtered calibration coefficients.



Residual RMS detects some bad quality spectra not flagged in L1C

What causes the remaining outliers?

Anomaly related to Metop-B manoeuvre on 20130807:

Met-office noticed a sudden increase in bias over Brazil when the manoeuvre occurred (plots from Nigel Atkinson)







Searching for signal in the residuals.



Good luck!

Normalised residual at 668 cm-1, 20091181 Asc. (spatial correlation 0.26)

Normalised residual at 667.5 cm-1, 20091181 Asc. (spatial correlation 0.35)





Residual spatial correlation (20110202)



Lets look at some PC scores and guess what we see



Band 2, PC score 1, Pixel 1 (20110202 12-24)



Band 2, PC score 2, Pixel 1 (20110202 12-24)

Band 2, PC score 24, Pixel 1 (20110202 12-24)



Band 2, PC score 21, Pixel 3 (20110202 12-24)



Some PC scores are very sensitive to spectral shift.

(measured by the variance of the PC score computed from spectrally shifted spectra divided by the original variance)



Black dots correspond to inhomogeneous fields of view \rightarrow non-uniform scene ILS effects

Do you think the upwelling radiance is correlated with IASI's cube corner direction?



Probably not, but several directions of the measurements are!

Same PC score, 4 different detectors



Band 2, PC score 24, Pixel 2 (20110202 12-24)



90 60° 30° 0 -30 30° -60 60 -90 00 180 240 300 0 60 120 180 -97 2603 -75.6469 -54.0335 -32,4201 -10.8067 10.8067 32,4201 54.0335 75.6469 97 2603

Band 2, PC score 24, Pixel 3 (20110202 12-24)

Band 2, PC score 24, Pixel 4 (20110202 12-24)



Evidence of instrument artefacts and inter-pixel differences \rightarrow identify the affected directions by plotting the means and standard deviations of the scores computed individually for each detector (normalised because of highly variable dynamic range of the scores)

>PC score standard deviation in each pixel (divided by average standard deviation)

> Mean of PC score in each pixel (minus average mean and divided by average standard deviation)

Spatial correlation of PC score in each pixel

>Inter EFOV spatial correlation of PC score







The signal and forward model subspaces

 $E_S \in \mathbb{R}^{m \times p}$ $E_F \in \mathbb{R}^{m \times p}$

Spanned by truncated sets of eigenvectors

Intersection of the two subspaces is empty → Compute principal angles between the two subspaces

$$E_S^{\ T} E_F = U S V^T$$
$$\widehat{E_S} = E_S U \qquad \widehat{E_F} = E_F V$$

 \hat{E}_{F} and \hat{E}_{S} are bi-orthogonal bases for the two subspaces.

The principal angles between the two subspaces are given by $\cos^{-1}(S_{ii})$ in ascending order

New bases for the signal and forward model spaces, in which similar directions are identified and ordered according to their degree of similarity







IASI data are not fully exploited

Often better results could be obtained simply by using reconstructed instead of raw radiances
 The problem is that people expect the same results and get disappointed when in fact they get better (but different) results



Ammonia signal near Lake Baikal, 18 April 2008 (Reconstructed radiances)





More cases close to and above the threshold than close to and below the threshold \rightarrow Number of cases classified as cloudy decreases after noise filtering.

Neural network retrievals taking subset of IASI channels as input

- make retrievals with raw radiances
- make retrievals with reconstructed radiances
- compare the two \rightarrow differences "too big" \rightarrow reconstructed radiances are bad (?!)

Should instead compare the two retrievals with an independent "truth"

> Due to the non-linearity of NN the best solution would be to train with reconstructed radiances

Using subset of reconstructed radiances instead of PC scores for retrieval or assimilation

The two cost functions are identical if

- 1. The forward model space is a subspace of the signal space, or alternatively, the reconstructed radiances are projected onto the forward model space.
- 2. The channel subset of reconstructed radiances is chosen such that the corresponding sub-matrix of the eigenvectors is invertible.
- + easy to reject "contaminated channels"
- + faster than using current PC forward models
- (-) need dense linear algebra for the observation error covariance, but really you always need it

"We want to get raw radiances because we want use our own PC's!"

-Good reasons to use own (forward model) PC's

-But does that mean that radiances should not be disseminated using another set of PC's?

To answer this question we look at the difference between the two scenarios:

- 1. Projecting raw radiances onto forward model space
- 2. Projecting raw radiances onto signal space and then onto forward model space
- Clearly if the forward model space was a subspace of the signal space there would be no difference.
- However, this is not the case in practise! (Forward model produce features never seen in any real measurements)

Two interpretations of this:

- a) Forward model errors
- b) Signal outside the signal space (i.e. reconstruction error)









Can PC's obtained from IASI-A spectra be used for dissemination of IASI-B spectra?



Yes. That's what we're doing and it works fine.

Ruminations on IASI-A vs. IASI-B



Eigenvectors E_A and E_B based on the covariance of a complete day of measurements from IASI-A and B



wavenumber (cm⁻¹)

 $\cos^{-1}(S_{ii})$ rank

http://www.eumetsat.int/website/home/Data/Products/Level1Data/

IASI Level 1 PCC Product Generation Specification, EUM.OPS-EPS.SPE.08.0199 IASI Level 1 PCC Product Format Specification, EUM.OPS-EPS.SPE.08.0195 IASI Principal Component Compression (IASI PCC) FAQ

Eigenvectors shared with EARS-IASI (NWPSAF collaboration(Nigel Atkinson, Fiona Smith))

Dissemination of Metop-A IASI PC scores on EUMETCast since 2010.08.05

IASI PCC Eigenvector files - Band 1,2 and 3 (HDF5) 1.3

EPS Product Validation Report: IASI L1 PCC PPF (Part 1), EUM/OPS-EPS/REP/10/0148

Metop-A IASI PC scores product operational since 2011.02.22

IASI PCC Eigenvector files - Band 1,2 and 3 (HDF5) 1.4

EPS Product Validation Report: IASI L1 PCC PPF (Part 2), EUM/OPS-EPS/REP/11/0036

Metop-B IASI PC scores product operational since 2013.03.12

Product validation report. IASI PCC for Metop-B, EUM/RSP/TEN/13/691073

Conclusions

• no serious compression without noise removal → danger of removing signal as well → however except for extremely rare cases detected by the residual RMS and 5 channels around 667.75 cm very little sign of this can be found in the residuals

- quality control of L1C spectra possible with the residual RMS
- decomposing spectra in the directions of the eigenvectors reveals small but undesired artefacts very clearly
- IASI measurements not exploited to their full potential, because users (of channel subsets) insist on keeping all the noise



The end



Why do we use separate PC's for each of the three IASI bands?

→ If one band (typically band 3) is of bad quality we can disseminate good PC scores from the other bands

+ Compression/decompression is faster

- About 40 extra PC scores are needed

Band	1	channel#	0	to	1996	(1997)
Band	2	channel#	1997	to	5115	(3119)
Band	3	channel#	5116	to	8460	(3345)

Band 1: 645.00 - 1144.00 cm-1 Band 2: 1144.25 - 1923.75 cm-1 Band 3: 1924.00 - 2760.00 cm-1

Experiments show that about 40 PC scores less would be needed to reach the same level of residual RMS as with band separation →



Noise normalisation

Diagonal noise normalisation matrix.

- noise does not get de-correlated ☺
- it works, but is suboptimal and should be changed (Nigel Atkinson, Fiona Hilton)

N equal to the matrix square root of the instrument noise covariance matrix.

- the correlated L1C noise gets normalised and de-correlated.
- equivalent to de-apodising prior to compression
- ullet ensures that same amount of noise is carried by all eigenvectors \odot



 $y = y_0 + \varepsilon_y$ Measured radiances = "true radiances" + noise

$$p = E^T N^{-1} (y - \bar{y})$$
 PC scores
 $\tilde{y} = NEp + \bar{y}$ Reconstructed radiances

 $A_y = NEE^T N^{-1}$ (short hand notation)

$$\tilde{y} = A_y(y - \bar{y}) + \bar{y} = A_y y + (I - A_y)\bar{y}$$

a projection (a linear transformation P from \mathbb{R}^{8461} to itself such that $P^2 = P$)

two orthogonal subspaces providing a unique decomposition of each spectrum

range(P)	Signal space
kernel(P)	Noise space



Practical issues

• Execution speed

Compress/reconstruct many spectra simultaneously, i.e. use matrixmatrix multiplications instead of matrix-vector multiplications

- $E^T N^{-1}$ and NE can be pre-computed No execution time penalty for using non-diagonal N / N^{-1}
 - Quantisation of PC scores

Dynamic range of PC scores decreases with the rank, most scores can be stored in one byte

• Use update formulas for covariance matrix when adding outlier spectra to the training set

Residual statistics, one full day (20100321)





(Noise Normalised) Residual standard deviation for each pixel, 20110202

Reconstructed BT (K) at 2380.5.5 cm-1, 20100516_D



