

# **Compression of AIRS data with principal components**

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**With acknowledgements to :**

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# Aims

1. To compare existing encoded AIRS products with model data
2. To investigate different methods of eigenvector creation

## Outline of talk

- i. Encoding data with principal components
- ii. Comparison with ECMWF model background
- iii. Initial assimilation trial with ECMWF model
- iv. Creation of PC sets: all-sky vs clear
- v. Test of reconstruction error



# Encoding of a spectra

- Given a set of spectral eigenvectors, arranged as columns of a matrix  $\mathbf{U}$ , an observation,  $obs$ , is coded into a vector of coefficients,  $c$ , by

$$c = \mathbf{U}^T obs$$

➤ Where  $T$  denotes the transpose of the matrix

- The reconstructed spectra,  $recon$ , is calculated from  $c$  by :

$$recon = \mathbf{U} c$$

- Spectral features that cannot be represented by the given eigenvectors will not be included in  $recon$ .
- These missing features can be summarised into a Reconstruction Error,  $RE$ , calculated for each spectra from :

$$RE = \sqrt{\frac{1}{n} \sum_{i=0}^n (obs_i - recon_i)^2}$$

➤ where  $n$  is the number of channels (after Goldberg *et al.* 2003)

➤ Subscript  $BT$ ,  $RE_{BT}$ , will be used when  $obs$  &  $recon$  are brightness temperatures

➤ Subscript  $R$ ,  $RE_R$ , will be used when  $obs$  &  $recon$  are in noise normalised radiance units



# Components of RE

$$RE = \sqrt{\frac{1}{n} \sum_{i=0}^n (obs_i - recon_i)^2}$$

Information that ends up in RE includes :

**1.  
Noise**

**Random noise in a particular channel not reconstructed - information from other channels**

**"de-noising"**

**2.**

**Spectral features due to the atmospheric state**

- i. Useful information, such as a structure not in U**
- ii. Information outside NWP model, ie aerosol**



## **ii. Comparison with ECMWF model background**



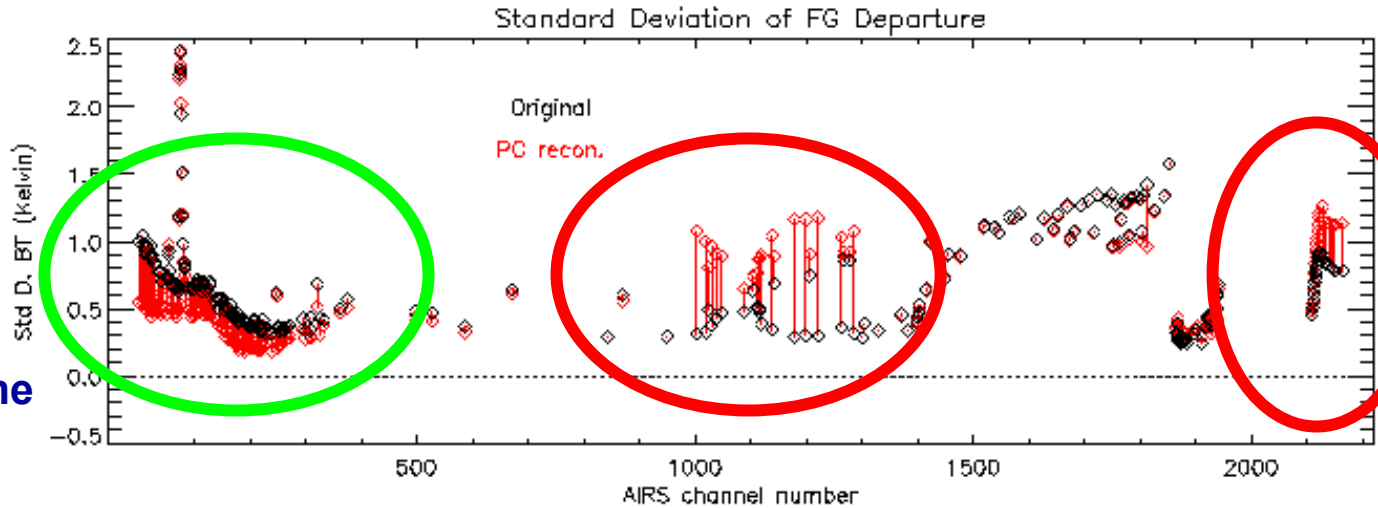
# Comparing NESDIS PC data with ECMWF model

- **“1524” Principal Component set from NESDIS :**
  - Used 1,524 channels
  - created from data over 1 day, 20<sup>th</sup> December 2002 (thinned to reduce the volume only)
  - PC coefficients calculated for every central AIRS view from alternate “golfballs” - “U1” dataset of 1 / 18<sup>th</sup> data
  - first 200 PC coefficients transmitted for reconstruction using eigenvector set
- **Comparison over 12 hour cycle with:**
  - background from ECMWF operational model (CY26R3) and
  - standard 324 channel data



# NESDIS PC departures from EC model

- Mean  $RE_{BT} \sim$  zero
- Some channels show higher  $RE_{BT}$  standard deviation (blue)

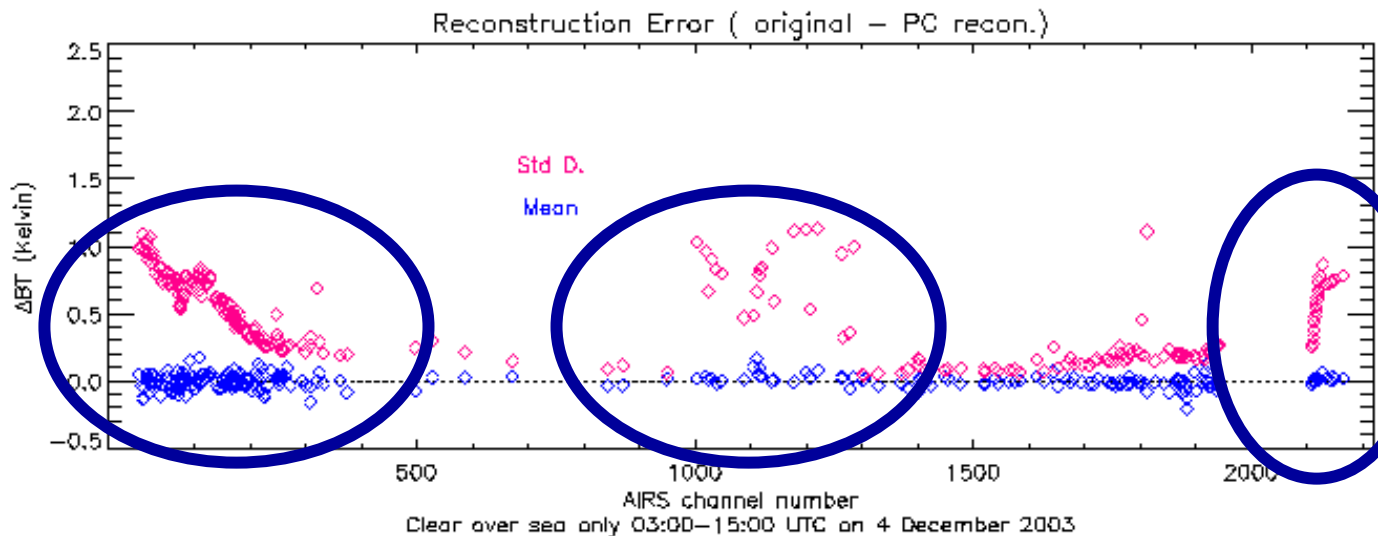


some also show increased std deviation from model (red)

“tropical” daytime feature ?

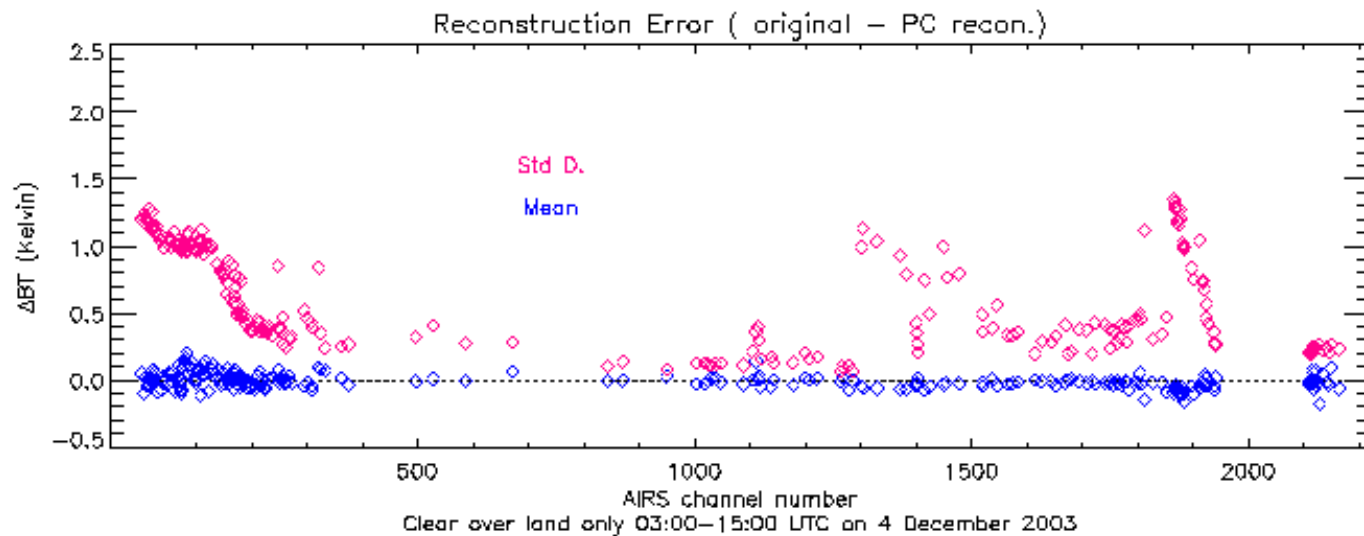
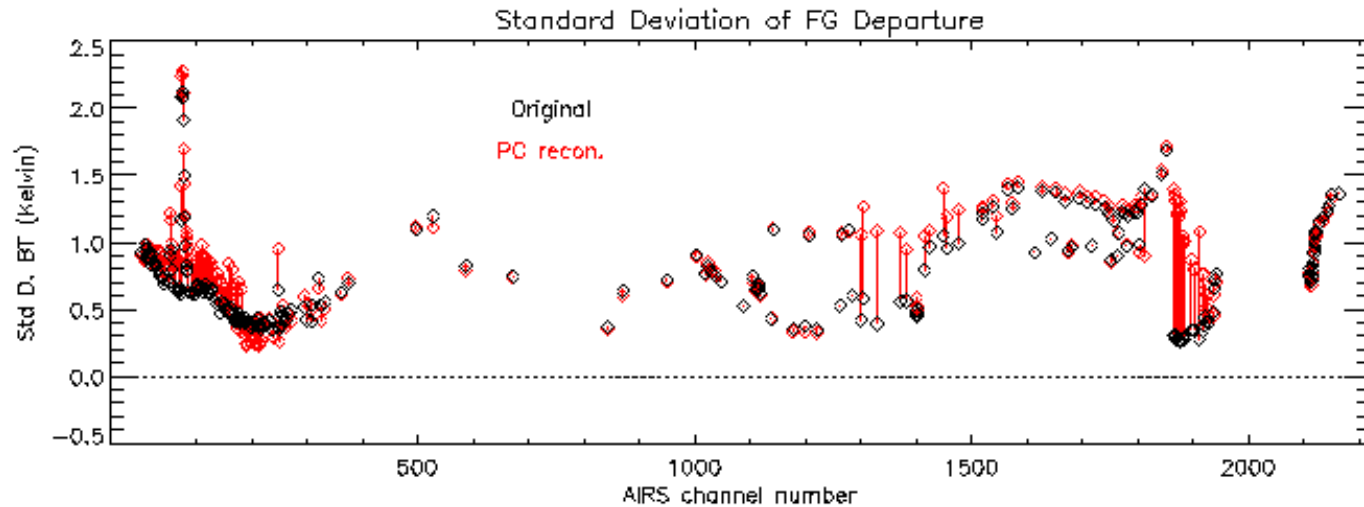
Others - high, stratospheric, channels - show decrease in std deviation from model (green)

“de-noised”



# Over land & at night ...

- At night increased SD went away
- Over land de-noising went away
- 12 hours data in January
  - still shows de-noising
  - Increased SD different





# **iii. Initial assimilation trial with ECMWF model**



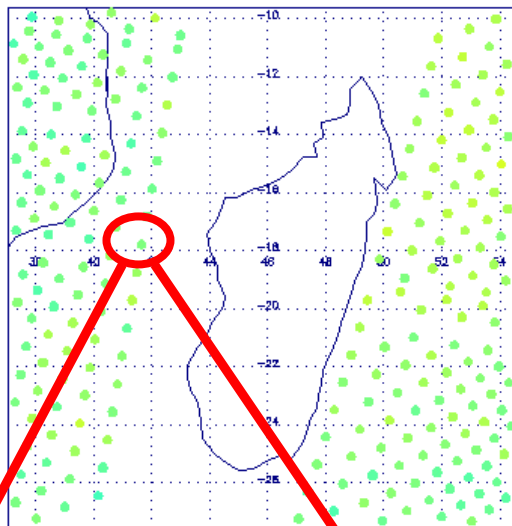
# Reconstructed Radiance assimilation

- **NESDIS “RR” data**
  - as the NESDIS “1524” PC data, except based on 2,047 channels PC’s
  - Delivered as BT’s over 322 channels
  - 200 coefficients used in reconstruction
- **Assimilated into EC operational model**
- **For a first attempt, error characteristics unchanged**
  - ... two following slides from Tony McNally

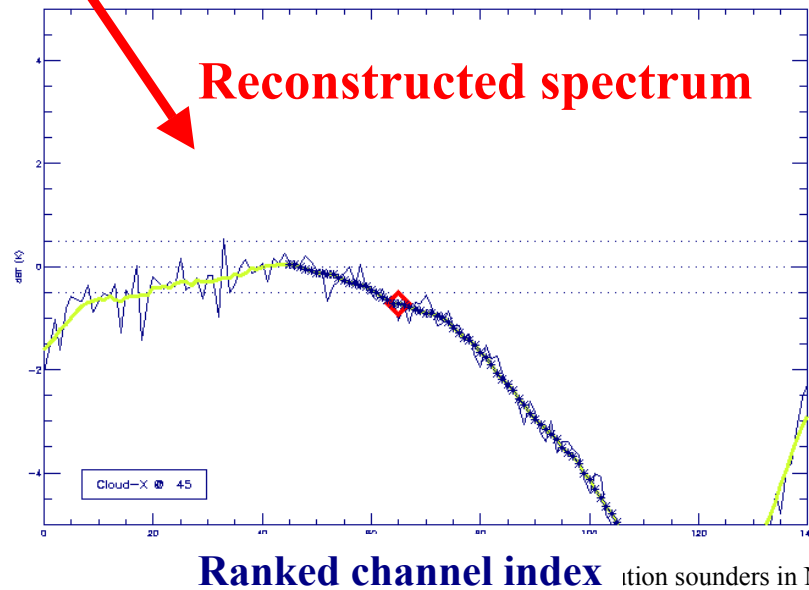
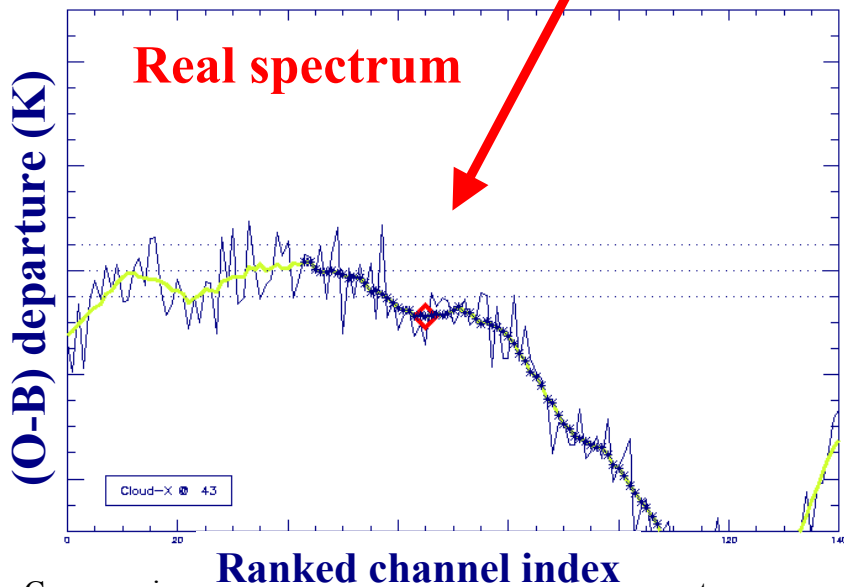


# De-noising with 200 NESDIS principal components

The PC de-noising influences the filtering used to reduce noise during the cloud detection process.  
(2 ranked channel change)



Channel to channel departure reduced

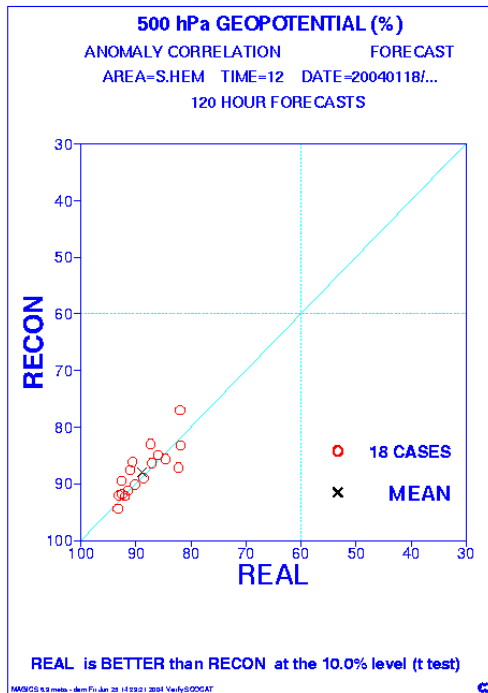
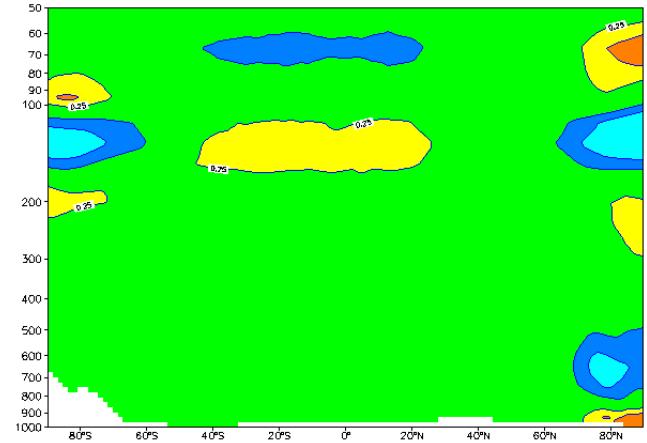


# Assimilation experiments with NESDIS PC reconstructed radiances

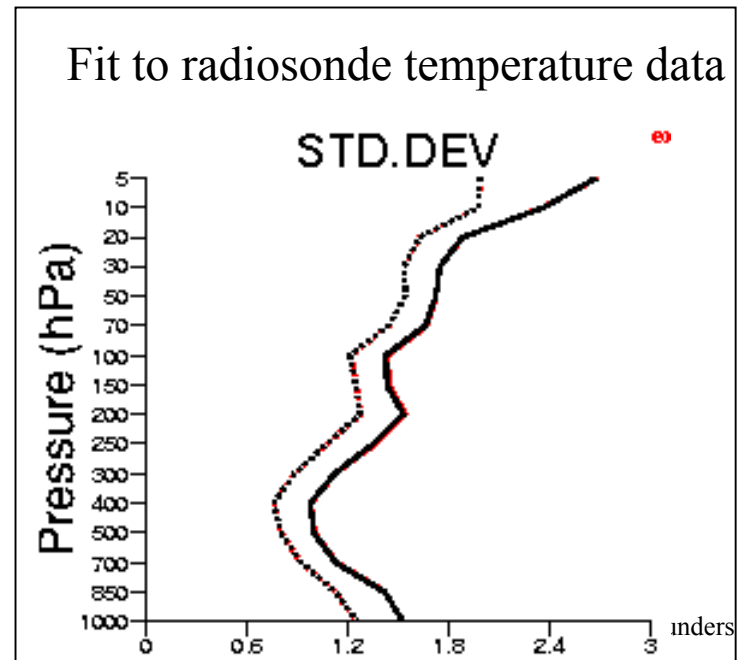
Slightly larger stratospheric analysis increments obtained with the reconstructed (de-noised) radiances

possible *organisation* of radiance signal ?

Zonal mean of RMS analysis increments



No impact on forecast quality or analysis fit to other data



ul components.

## **iv. Creation of PC sets: all-sky vs clear**



# Motivations for study

- **De-noising to suit NWP application**
- **Investigate different “training” sets for PC**
  - Can PC based only on clear views better reconstruct clear views ?
  - Clear based PC allow addition of cloud signal eigenvectors :
    - **Gather residual spectra from several cloudy views**  
$$residual = obs - \mathbf{U}c$$
    - **Singular value decomposition to create eigenvectors**
    - **Concatenated into a new  $\mathbf{U}$  with new coefficients added to end of  $c$**
    - **Can be repeated for other scenes ... .**



# Principal Component test sets

- **Two sets from data in July 2003**
  - “All” set
    - one day (15<sup>th</sup>), all views, all angles, land & sea, 1/9<sup>th</sup> thinned ( 324,000 spectra )
  - “Clears” set
    - one day (16<sup>th</sup>), over sea, clear at AIRS Level 2, unthinned ( gave 85,000 spectra, ~3% *clear* )
    - Added further high latitude ( $> 40^\circ$ ) clears from 15<sup>th</sup>
- **For both**
  - 2,107 channels used
    - channels valid for AIRS Level 2
  - Radiance data, noise normalised
    - instrument noise taken from channel properties file
  - calculate spectral covariance matrix and derive eigenvectors from that



# PC sets' variance

## ➤ Diagonal of covariance matrices

○ In noise normalised radiance

○ Black - "All"

○ Red - "Clears"

## ➤ Spectral Bands

1. high peak LW CO<sub>2</sub>

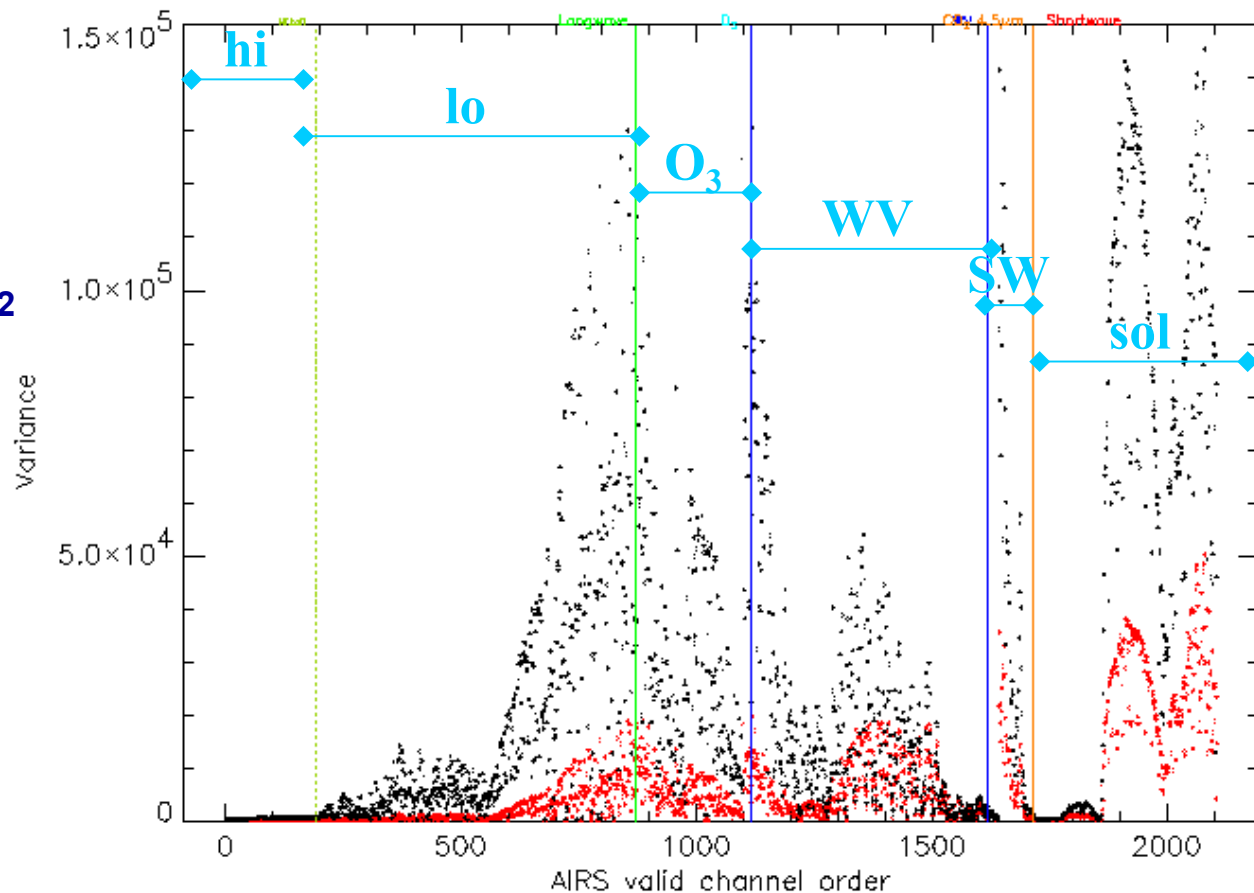
2. low peak LW CO<sub>2</sub>

3. ozone

4. water vapour

5. SW CO<sub>2</sub>

6. solar end





## **v. Test of reconstruction error**



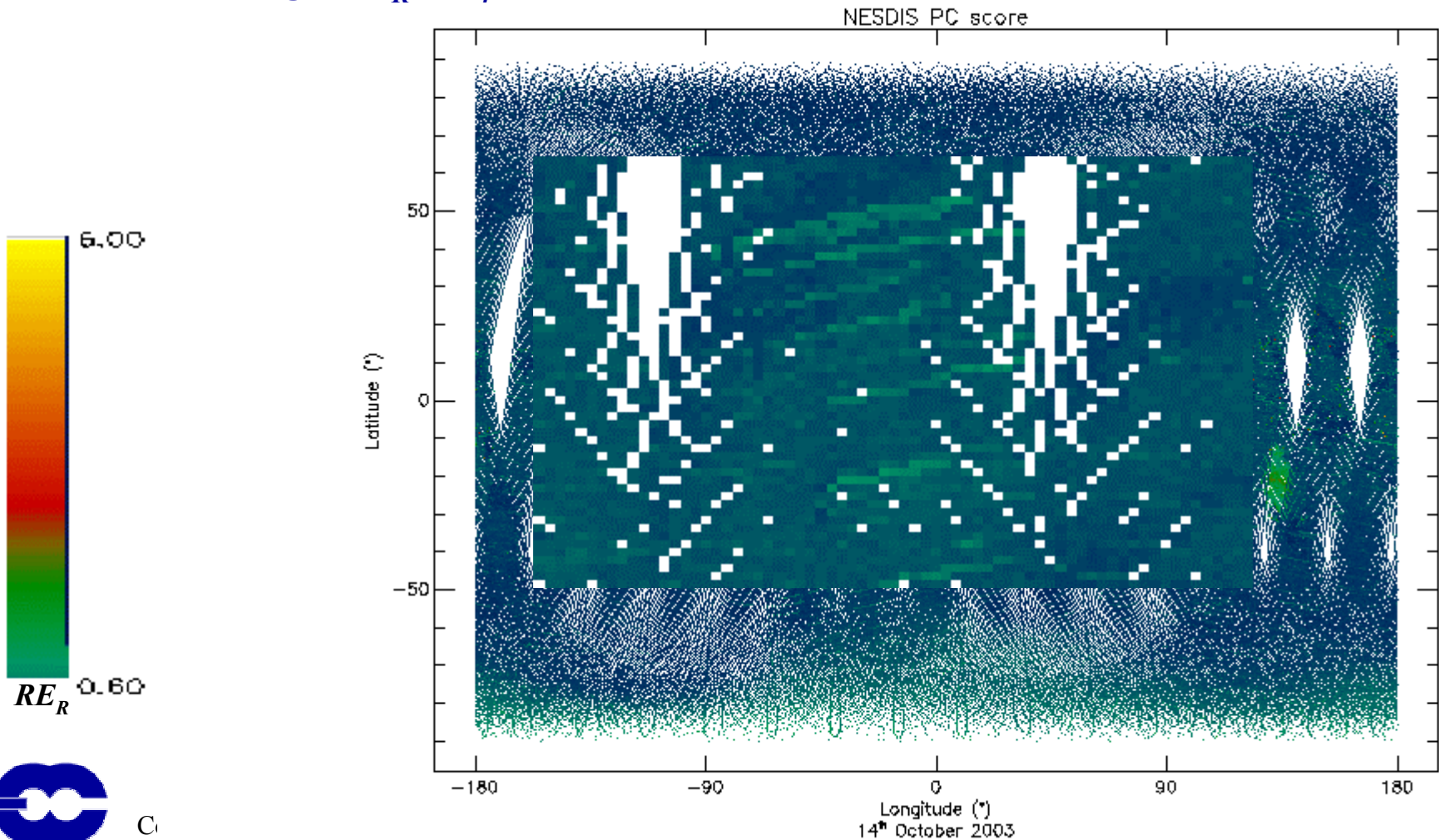
# Full spectra comparison data

- **Illustrative data from 1 in nine spectra from one day - 14<sup>th</sup> October 2003**
  - (processed granules 1 to 188)
- **“All” view based PC compared with “Clears” PC**
- **Maps of  $RE_R$**



# Mapped $RE_R$

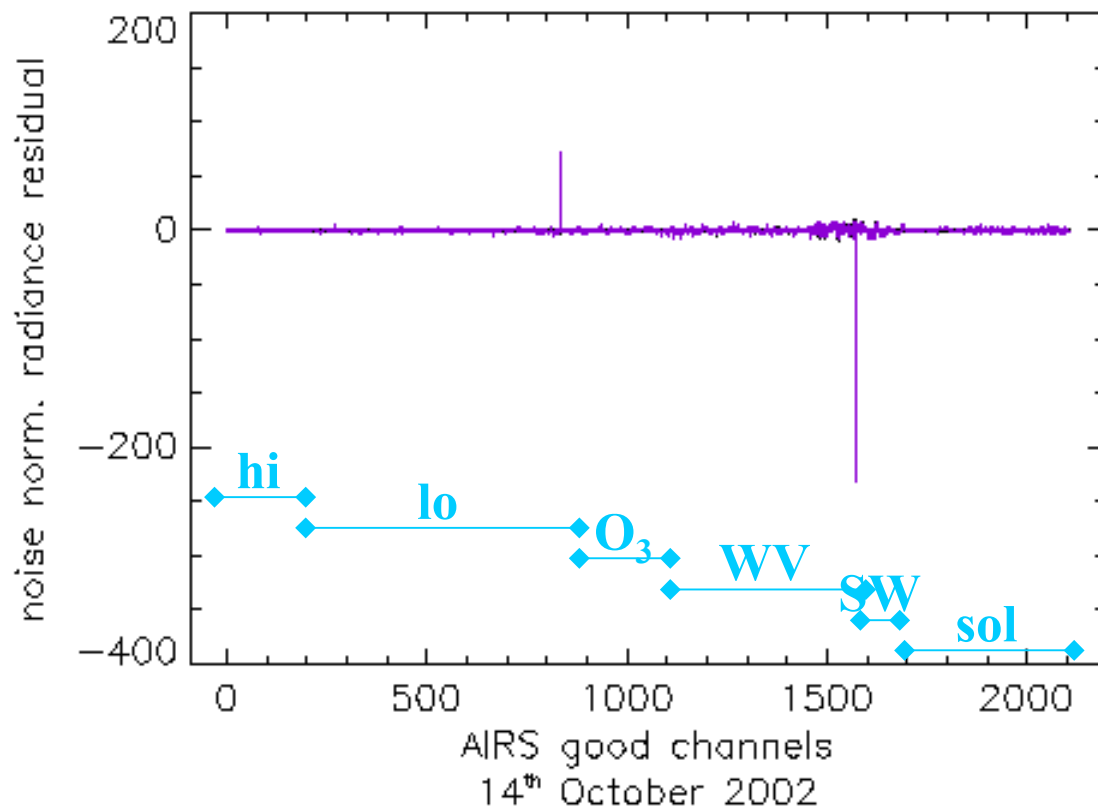
- clear views from Level 2 - mainly between 40°N to 40°S
- 63 % of clear spectra have lower  $RE_R$  with “Clears” PC
- for average  $RE_R$ , 4% lower than NESDIS PC, 26% lower than NESDIS PC



# Results - time consistency

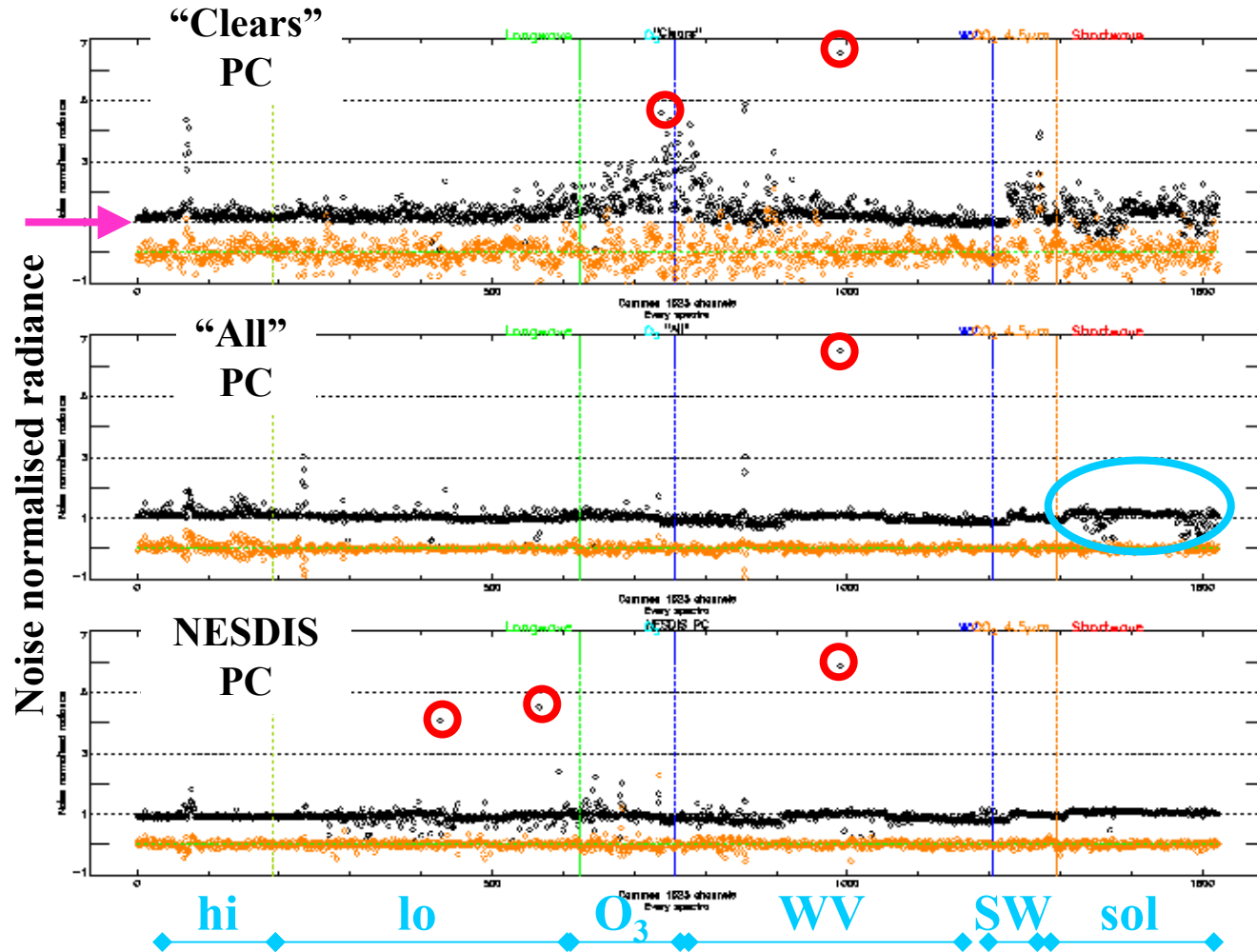
## ■ *RE* compromised by channel changes

- Plot shows  $RE_r$  from October 2002 using “All” & “Clears” PC sets (derived from July 2003 data)
- Spikes from channels :  
957 ( $982.0 \text{ cm}^{-1}$ ) &  
1,791 ( $1561.6 \text{ cm}^{-1}$ )  
(pop & noise)



# All views

- Analysis per channel of radiance reconstruction error
- Std Deviation in black and mean in orange
  - Over the 1,523 common channels in “All”, “Clears” & NESDIS PC
  - Data from 14th October 2003 from 254,800 spectra

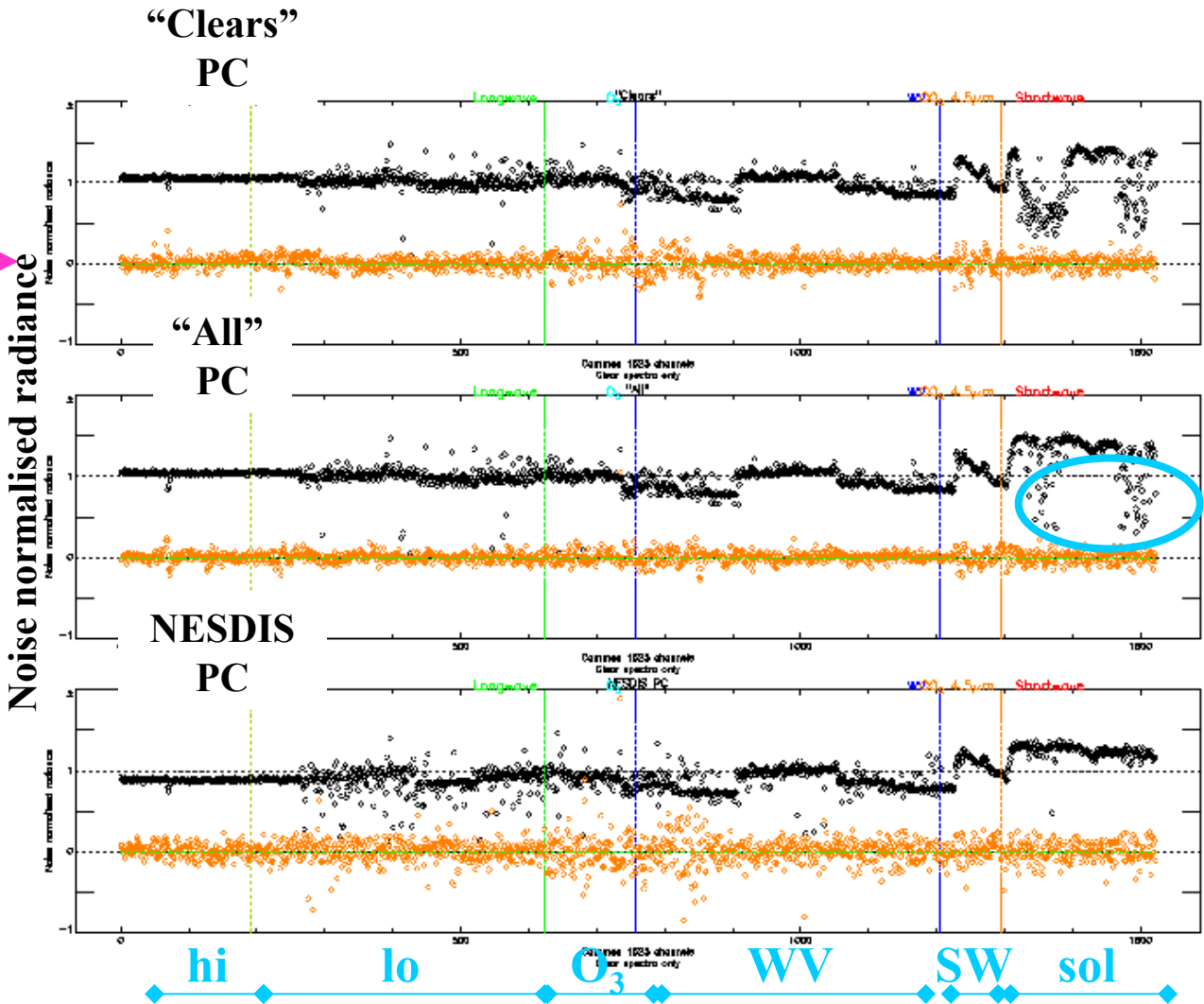


# Clear views

- **Less noise in “clears”**

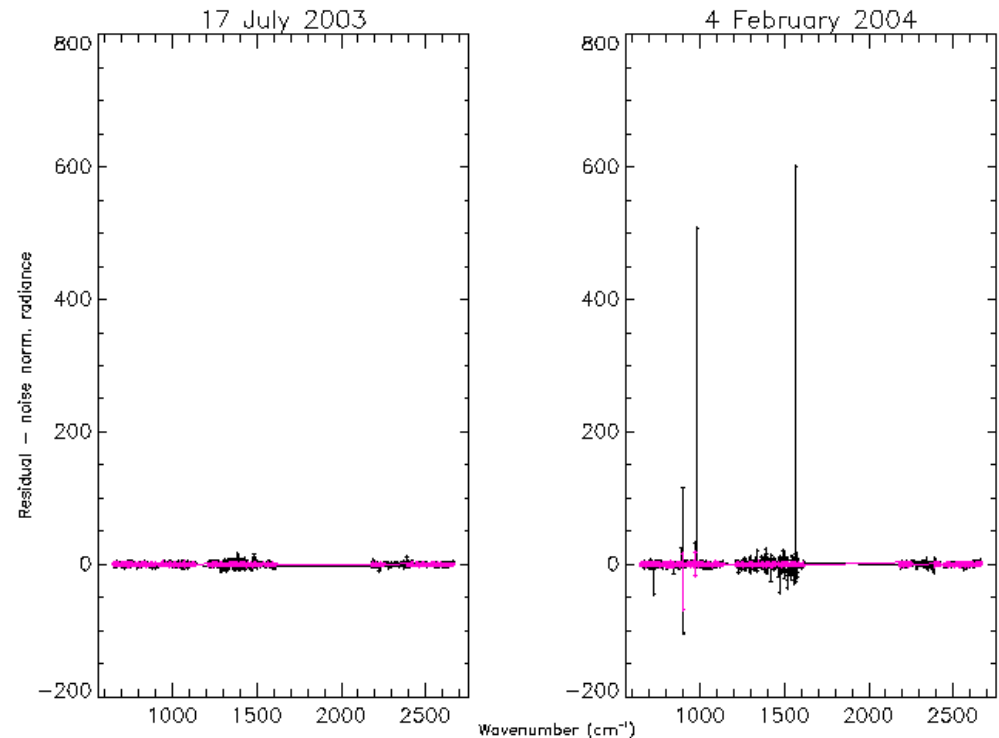
- NESDIS Std Dev. less for high LW CO<sub>2</sub>

- (outliers still present - offscale)



# Solar shutdown changes

- Several detectors not the same on power up
- Show up in  $RE_r$
- Comparison of July '03 and February '04
  - Spikes in February data for channels :  
756 (899.3  $\text{cm}^{-1}$ ),  
765 (902.4  $\text{cm}^{-1}$ ),  
957 (982.0  $\text{cm}^{-1}$ ) &  
1,802 (1569.3  $\text{cm}^{-1}$ ) none in the 324 operational channels
  - Operational channels changed were :  
318 (741.3  $\text{cm}^{-1}$ ), 1,883 (2197.9  $\text{cm}^{-1}$ ) & 1,884 (2198.8  $\text{cm}^{-1}$ ) (FG dep. > 3 x Std Err.)



# Conclusions

## ■ De-noising

- NESDIS PC compared with model first guess
  - Seen for stratospheric channels
  - Other channels the same or noisier
- NESDIS RR assimilated, as above plus
  - Increased stratospheric increments
- Need not “de-noise” all channels if residual available
- Monitoring needs RE and FG departure

## ■ “Clears” PC

- Improvement found for majority of clear views
  - ✗ It was only a small majority
  - ? Improvement big enough ?
  - ? Tests so far only over limited number of clears from Level 2
  - ? Too many bad channels





# Conclusions (continued)

- **In particular for AIRS :**
  - channels that go bad are a long term problem
  - line by line offset significant when looking for the one clear view for NWP



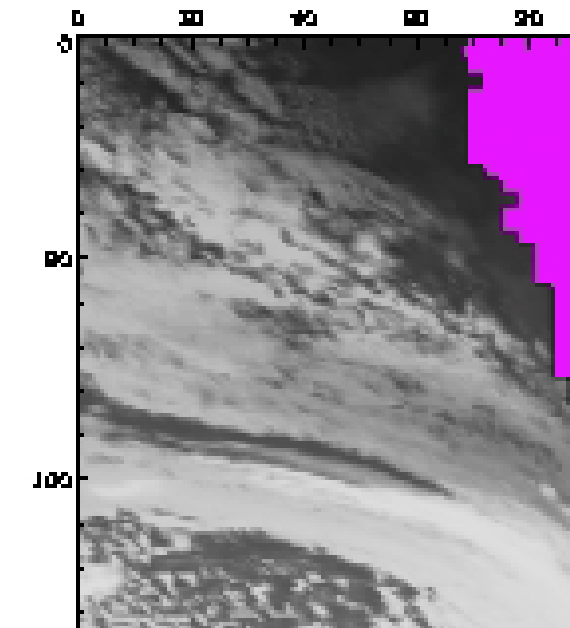
**?. but ....**



# Better clear PC set

- PC set similar to “clear-sky”
- But 75 % of spectra not normalised by noise
- Decreased mean
- Overall increase in variance, less at shortwave

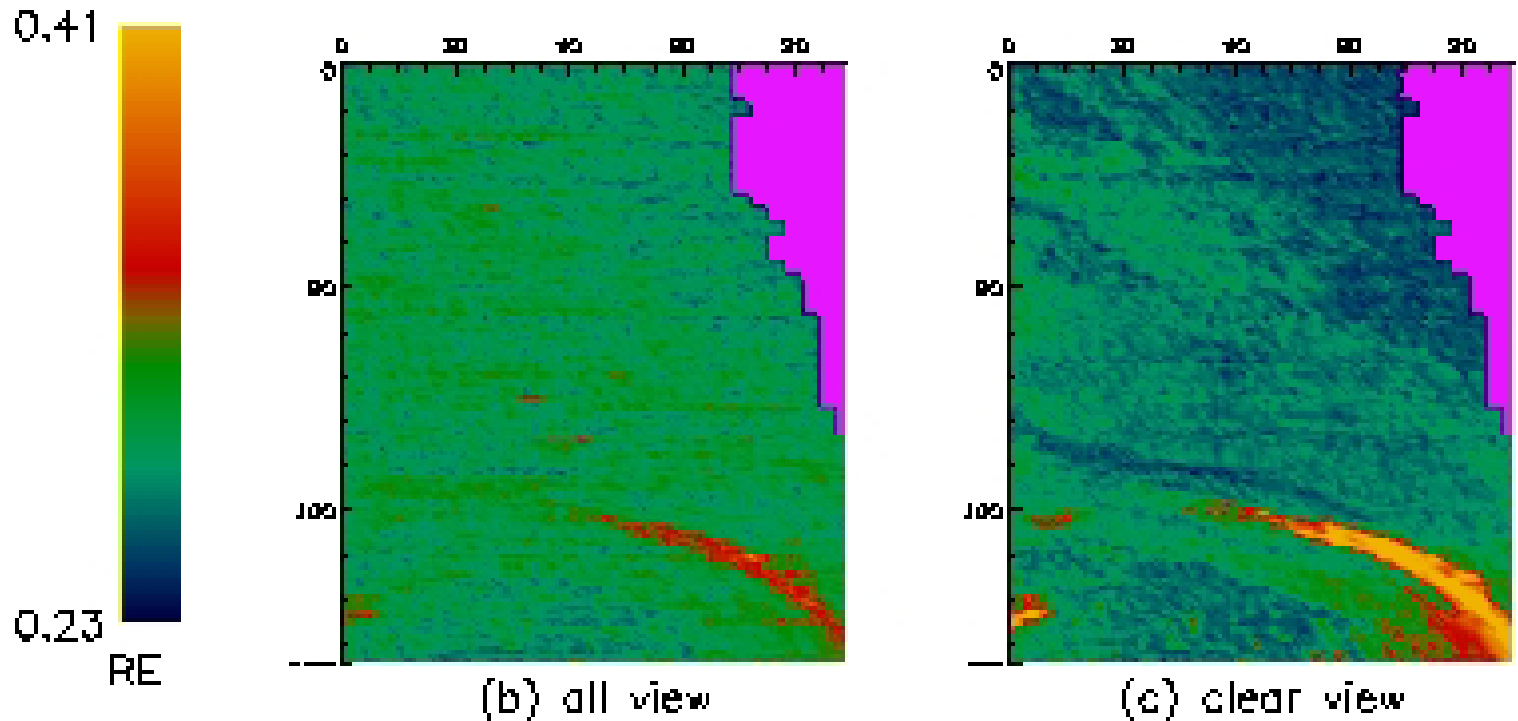
**Example: Western  
Australia (purple),  
Granule 180, 15<sup>th</sup>  
October 2003**



(a)  $\lambda = 10.91 \mu\text{m}$



# Lower $RE_{BT}$



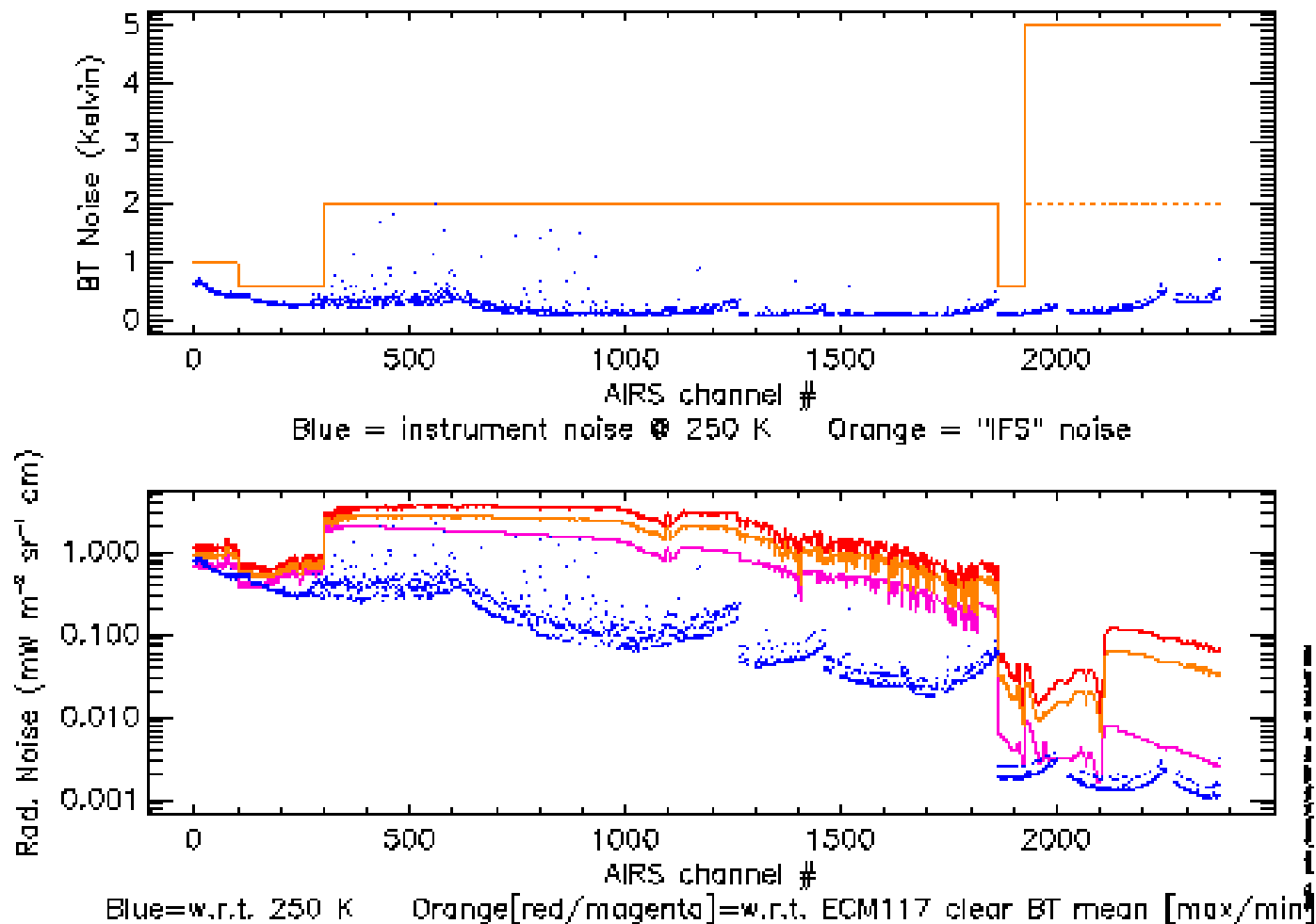
- **83% of clears with lower  $RE_{BT}$  than NESDIS PC set**



**extra . . . . .**



# Noise



# Lines across $RE_r$

- **Due to offset calibration at end of each line**
  - details of “striping” in Steve Gaiser’s presentation March ‘04 Science Team meeting
  - Overall bias zero - but adds 5% noise overall
- **Line to line variability significant**
- **Example .....**

