## IFS DOCUMENTATION

# Part III: DYNAMICS AND NUMERICAL PROCEDURES (CY25R1) 

(Operational implementation 9 April 2002)

Edited by Peter W. White
(Text written and updated by members of the ECMWF Research Department)

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## Part III: Dynamics and numerical procedures

## CHAPTER 1 Introduction

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1.1 Overview

### 1.1 OVERVIEW

Since the original demonstration of the efficiency advantage of the semi-Lagrangian semi-implicit method over a decade ago by André Robert, this numerical integration scheme is being used in an increasing range of atmospheric models. Most of the applications have been in grid-point models. Shallow-water-equations studies have included three-time-level versions by Robert (1981, 1982) and Staniforth and Temperton (1986), and two-time-level schemes by Temperton and Staniforth (1987), Purser and Leslie (1988), McDonald and Bates (1989), and Côté and Staniforth (1990). There also have been various applications in baroclinic grid-point models. Three-time-level sigma-coordinate versions have been presented by Robert et al. (1985) and Tanguay et al. (1989), and the extension of the three-time-level approach to a non-hydrostatic coordinate has been demonstrated by Tanguay et al. (1990) . Bates and McDonald (1982), McDonald (1986), Leslie and Purser (1991), McDonald and Haugen (1992), and Bates et al. (1993) have developed two-time-level sigma-coordinate schemes, McDonald and Haugen (1993) have presented the two-time-level extension to a hybrid vertical coordinate, and Golding (1992) has applied a split two-time-level semi-Lagrangian scheme in a non-hydrostatic model.

For spectral models, a semi-Lagrangian semi-implicit shallow-water equation model was presented by Ritchie (1988) for a three-time-level version, and adapted by Côté and Staniforth (1988) for a two-time-level scheme. Baroclinic three-time-level spectral model formulations have been demonstrated by Ritchie (1991) for operational numerical weather prediction in a sigma-coordinate model, and recently by Williamson and Olson (1994) for climate simulations with a hybrid coordinate model.

In a broader context, the semi-Lagrangian scheme, as incorporated in spectral numerical weather prediction models, may be considered as an economical variant of the spectral Lagrange-Galerkin method (Süli and Ware, 1991).

Experience at ECMWF (Simmons et al., 1989) suggests that the accuracy of medium-range forecasts has steadily improved with increasing resolution. Consequently, in its four-year plan for the period 1989-1992, ECMWF proposed development of a high-resolution version of its forecast model. A target resolution of a spectral representation with a triangular truncation of 213 waves in the horizontal and 31 levels in the vertical (T213/L31) was set, entailing a doubling of the horizontal resolution and an approximate doubling of the vertical resolution in the troposphere compared to the T106/L19 configuration that was operational at the time (Simmons et al., 1989). In view of the anticipated computer resources, it was clear that major efficiency gains would be necessary in order to attain this objective. These gains have been provided by the introduction of the semi-Lagrangian treatment of advection permitting a substantial increase in the size of the time-step, the use of a reduced Gaussian grid giving a further advantage of about $25 \%$, the introduction of economies in the Legendre transforms, and improvements to the model's basic architecture.

The layout for the remainder of the document is as follows. In Chapter 2 'Basic equations and discretization' we present the reformulation of the Eulerian model in order to transform the vorticity-divergence formulation into a
momentum-equation version in preparation for a subsequent semi-Lagrangian vector treatment of the equations of motion. The vertical discretization of the ECMWF hybrid coordinate on a staggered grid is also considered. The semi-Lagrangian treatment is discussed in some detail in Chapter 3 'Semi-Lagrangian formulation', including the adaptation to accommodate the reduced Gaussian grid. Several important computational details relevant for efficient execution of the high resolution model on a modestly parallel supercomputer are discussed in Chapter 4 ‘Computational details’.

## Part III: Dynamics and numerical procedures

## CHAPTER 2 Basic equations and discretization

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### 2.1 EULERIAN REFORMULATION OF THE CONTINUOUS EQUATIONS

Following Ritchie $(1988,1991)$, the first step in developing a semi-Lagrangian version of the ECMWF spectral model was to convert the existing Eulerian $\zeta-\mathrm{D}$ (vorticity-divergence) model to a $\mathrm{U}-\mathrm{V}$ formulation, where U and V are the wind images defined by $\mathrm{U}=\mathrm{u} \cos \theta, \mathrm{V}=\mathrm{V} \cos \theta(\mathrm{u}$ and V are the components of the horizontal wind in spherical coordinates, and $\theta$ is latitude). In this section we describe the Eulerian $\mathrm{U}-\mathrm{V}$ model.

First we set out the continuous equations in $(\lambda, \theta, \eta)$ coordinates, where $\lambda$ is longitude and $\eta$ is the hybrid vertical coordinate introduced by Simmons and Burridge (1981); thus $\eta\left(p, p_{\text {surf }}\right)$ is a monotonic function of the pressure P , and also depends on the surface pressure $\mathrm{p}_{\text {surf }}$ in such a way that

$$
\eta\left(0, p_{\text {surf }}\right)=0 \text { and } \eta\left(p_{\text {surf }}, p_{\text {surf }}\right)=1
$$

The momentum equations are

$$
\begin{align*}
& \frac{\partial U}{\partial t}+\frac{1}{a \cos ^{2} \theta}\left\{U \frac{\partial U}{\partial \lambda}+v \cos \theta \frac{\partial U}{\partial \theta}\right\}+\dot{\eta} \frac{\partial U}{\partial \eta} \\
& (-f v)+\frac{1}{a}\left\{\frac{\partial \phi}{\partial \lambda}+R_{\text {dry }} \mathrm{T}_{\mathrm{v}} \frac{\partial}{\partial \lambda}(\ln \mathrm{p})\right\}=\mathrm{P}_{\mathrm{u}}+\mathrm{K}_{\mathrm{U}} \tag{2.1}
\end{align*}
$$

$$
\begin{align*}
\frac{\partial V}{\partial t}+\frac{1}{a^{2} \cos ^{2} \theta}\left\{U \frac{\partial V}{\partial \lambda}\right. & \left.+V \cos \theta \frac{\partial V}{\partial \theta}+\sin \theta\left(U^{2}+V^{2}\right)\right\}+\dot{\eta} \frac{\partial V}{\partial \eta} \\
& +f U+\frac{\cos \theta}{a}\left\{\frac{\partial \phi}{\partial \theta}+R_{\text {dry }} T_{v} \frac{\partial}{\partial \theta}(\ln \mathrm{p})\right\}=P_{\mathrm{V}}+K_{\mathrm{V}} \tag{2.2}
\end{align*}
$$

where $a$ is the radius of the earth, $\dot{\eta}$ is the $\eta$-coordinate vertical velocity $(\dot{\eta}=d \eta / d t), \phi$ is geopotential, $R_{\text {dry }}$ is the gas constant for dry air, and $T_{v}$ is the virtual temperature defined by

$$
\mathrm{T}_{\mathrm{v}}=\mathrm{T}\left[1+\left\{\mathrm{R}_{\text {vap }} /\left(\mathrm{R}_{\text {dry }}-1\right)\right\} \mathrm{q}\right]
$$

where $T$ is temperature, $q$ is specific humidity and $R_{\text {vap }}$ is the gas constant for water vapour. $P_{U}$ and $P_{V}$ represent the contributions of the parametrized physical processes, while $K_{U}$ and $K_{V}$ are the horizontal diffusion terms.

The thermodynamic equation is

$$
\begin{equation*}
\frac{\partial T}{\partial t}+\frac{1}{\operatorname{acos}^{2} \theta}\left\{U \frac{\partial T}{\partial \theta}+V \cos \theta \frac{\partial T}{\partial \theta}\right\}+\dot{\eta} \frac{\partial T}{\partial \eta}-\frac{\kappa T_{v} \omega}{(1+(\delta-1) q) p}=P_{T}+K_{T} \tag{2.3}
\end{equation*}
$$

where $\kappa=R_{\text {dry }} / c_{p_{\text {dry }}}$ ( $C_{p_{\text {dry }}}$ is the specific heat of dry air at constant pressure), $\omega$ is the $\rho$-coordinate vertical velocity $(\omega=d p / d t)$, and $\delta=C_{p_{\text {vap }}} / C_{p_{\text {dry }}}$ ( $C_{p_{\text {vap }}}$ is the specific heat of water vapour at constant pressure).
The moisture equation is

$$
\begin{equation*}
\frac{\partial q}{\partial t}=\frac{1}{a^{2} \cos ^{2} \theta}\left\{U \frac{\partial q}{\partial \lambda}+V \cos \theta \frac{\partial q}{\partial \theta}\right\}=\eta \frac{\partial q}{\partial \eta}=P_{q}+K_{q} \tag{2.4}
\end{equation*}
$$

In (2.2) and (2.3), $\mathrm{P}_{\mathrm{T}}$ and $\mathrm{P}_{\mathrm{q}}$ represent the contributions of the parametrized physical processes, while $\mathrm{K}_{\mathrm{T}}$ and $\mathrm{K}_{\mathrm{q}}$ are the horizontal diffusion terms.

The continuity equation is

$$
\begin{equation*}
\frac{\partial}{\partial t}\left(\frac{\partial p}{\partial \eta}\right)+\nabla \cdot\left(\mathbf{v}_{H} \frac{\partial p}{\partial \eta}\right)+\frac{\partial}{\partial \eta}\left(\dot{\eta} \frac{\partial p}{\partial \eta}\right)=0 \tag{2.5}
\end{equation*}
$$

where $\nabla$ is the horizontal gradient operator in spherical coordinates and $\mathbf{v}_{H}=(u, v)$ is the horizontal wind. The geopotential $\phi$ which appears in (2.1) and (2.2) is defined by the hydrostatic equation

$$
\begin{equation*}
\frac{\partial \phi}{\partial \eta}=\frac{R_{\text {dry }} T_{v}}{p} \frac{\partial p}{\partial \eta} \tag{2.6}
\end{equation*}
$$

while the vertical velocity $\omega$ in (2.3) is given by

$$
\begin{equation*}
\omega=-\int_{0}^{\eta} \nabla \cdot\left(\mathbf{v}_{\mathrm{H}} \frac{\partial \mathrm{p}}{\partial \eta}\right) \mathrm{d} \eta+\mathbf{v}_{\mathrm{H}} \cdot \nabla \mathrm{p} \tag{2.7}
\end{equation*}
$$

Expressions for the rate of change of surface pressure, and for the vertical velocity $\dot{\eta}$, are obtained by integrating (2.5), using the boundary conditions $\dot{\eta}=0$ at $\eta=0$ and at $\eta=1$

$$
\begin{align*}
& \frac{\partial p_{\text {surf }}}{\partial \mathrm{t}}=-\int_{0}^{1} \nabla \cdot\left(\mathbf{v}_{\mathrm{H}} \frac{\partial \mathrm{p}}{\partial \eta}\right) \mathrm{d} \eta  \tag{2.8}\\
& \dot{\eta} \frac{\partial \mathrm{p}}{\partial \eta}=-\frac{\partial p}{\partial \eta}-\int_{0}^{\eta} \nabla \cdot\left(\mathbf{v}_{\mathrm{H}} \frac{\partial \mathrm{p}}{\partial \eta}\right) \mathrm{d} \eta \tag{2.9}
\end{align*}
$$

Since we use $\ln \left(\mathrm{p}_{\text {surf }}\right)$ rather than $\mathrm{p}_{\text {surf }}$ as the surface pressure variable, it is convenient to rewrite (2.8) as

$$
\begin{equation*}
\frac{\partial}{\partial t}\left(\ln p_{\text {surf }}\right)=-\frac{1}{p_{s}} \int_{0}^{1} \nabla \cdot\left(\mathbf{v}_{H} \frac{\partial p}{\partial \eta}\right) \mathrm{d} \eta \tag{2.10}
\end{equation*}
$$

### 2.2 DISCRETIZATION

### 2.2.1 Vertical discretization

To represent the vertical variation of the dependent variables $\mathrm{U}, \mathrm{V}, \mathrm{T}$ and q , the atmosphere is divided into NLEV layers. These layers are defined by the pressures at the interfaces between them (the 'half-levels'), and these pressures are given by

$$
\begin{equation*}
\mathrm{p}_{\mathrm{k}+1 / 2}=\mathrm{A}_{\mathrm{k}+1 / 2}+\mathrm{B}_{\mathrm{k}+1 / 2} \mathrm{p}_{\mathrm{surf}} \tag{2.11}
\end{equation*}
$$

for $0 \leq k \leq N L E V$. The $A_{k+1 / 2}$ and $B_{k+1 / 2}$ are constants whose values effectively define the vertical coordinate and $\mathrm{p}_{\text {surf }}$ is the surface pressure field.

The values of the $A_{k+1 / 2}$ and $B_{k+1 / 2}$ for all $0 \leq k \leq N L E V$ are stored in the GRIB header of all fields archived on model levels to allow the reconstruction of the 'full-level' pressure $p_{k}$ associated with each model level (middle of layer) from $p_{k}=\frac{1}{2}\left(p_{k-1 / 2}+p_{k+1 / 2}\right)(1 \leq k \leq N L E V)$ by using (2.11) and the surface pressure field.
The prognostic variables are represented by their values at 'full-level' pressures $p_{k}$. Values for $p_{k}$ are not explicitly required by the model's vertical finite-difference scheme, which is described below.

The discrete analogue of the surface pressure tendency equation (2.10) is

$$
\begin{equation*}
\frac{\partial}{\partial \mathrm{t}}\left(\ln \mathrm{p}_{\text {surf }}\right)=-\frac{1}{\mathrm{p}_{\text {surf }}} \sum_{\mathrm{k}=1}^{\mathrm{NLEV}} \nabla \cdot\left(\mathbf{v}_{\mathrm{H}} \Delta \mathrm{p}_{\mathrm{k}}\right) \tag{2.12}
\end{equation*}
$$

where

$$
\begin{equation*}
\Delta \mathrm{p}_{\mathrm{k}}=\mathrm{p}_{\mathrm{k}+1 / 2}-\mathrm{p}_{\mathrm{k}-1 / 2} \tag{2.13}
\end{equation*}
$$

From (2.11) we obtain

$$
\begin{equation*}
\frac{\partial}{\partial t}\left(\ln \mathrm{p}_{\text {surf }}\right)=-\sum_{\mathrm{k}=1}^{\mathrm{NLEV}}\left\{\frac{1}{\mathrm{p}_{\text {surf }}} \mathrm{D}_{\mathrm{k}} \Delta \mathrm{p}_{\mathrm{k}}+\left(\mathrm{v}_{\mathrm{k}} \cdot \nabla \ln \mathrm{p}_{\text {surf }}\right) \Delta \mathrm{B}_{\mathrm{k}}\right\} \tag{2.14}
\end{equation*}
$$

where $D_{k}$ is the divergence at level $k$,

$$
\begin{equation*}
\mathrm{D}_{\mathrm{k}}=\frac{1}{\operatorname{acos}^{2} \theta}\left(\frac{\partial \mathrm{U}_{\mathrm{k}}}{\partial \lambda}+\cos \theta \frac{\partial \mathrm{V}_{\mathrm{k}}}{\partial \theta}\right) \tag{2.15}
\end{equation*}
$$

and

$$
\begin{equation*}
\Delta B_{k}=B_{k+1 / 2}-B_{k-1 / 2} \tag{2.16}
\end{equation*}
$$

The discrete analogue of (2.9) is

$$
\begin{equation*}
\left(\dot{\eta} \frac{\partial \mathrm{p}}{\partial \eta}\right)_{k+1 / 2}=-\frac{\partial p_{k+1 / 2}}{\partial t}-\sum_{j=1}^{k} \nabla \cdot\left(\mathbf{v}_{j} \Delta p_{j}\right) \tag{2.17}
\end{equation*}
$$

and from (2.11) we obtain

$$
\begin{equation*}
\left(\dot{\eta} \frac{\partial \mathrm{p}}{\partial \eta}\right)_{k+1 / 2}=-p_{\text {surf }}\left[B_{k+(1 / 2)} \frac{\partial}{\partial t}\left(\ln p_{\text {surf }}\right)+\sum_{j=1}^{k}\left\{\frac{1}{p_{\text {surf }}} D_{j} \Delta p_{j}+\left(v_{j} \cdot \nabla \ln p_{\text {surf }}\right) \Delta B_{j}\right\}\right] \tag{2.18}
\end{equation*}
$$

where $\partial / \partial t\left(\ln p_{\text {surf }}\right)$ is given by (2.14).
Vertical advection of a variable X is now given by

$$
\begin{equation*}
\left(\dot{\eta} \frac{\partial x}{\partial \eta}\right)_{k}=\frac{1}{2 \Delta p_{k}}\left\{\left(\dot{\eta} \frac{\partial p}{\partial \eta}\right)_{k+1 / 2}\left(X_{k+1}-X_{k}\right)+\left(\dot{\eta} \frac{\partial p}{\partial \eta}\right)_{k-1 / 2}\left(X_{k}-X_{k-1}\right)\right\} \tag{2.19}
\end{equation*}
$$

The discrete analogue of the hydrostatic equation (2.6) is

$$
\begin{equation*}
\phi_{\mathrm{k}+1 / 2}-\phi_{\mathrm{k}-1 / 2}=-\mathrm{R}_{\mathrm{dry}}\left(\mathrm{~T}_{\mathrm{v}}\right)_{\mathrm{k}} \ln \frac{\mathrm{p}_{\mathrm{k}+1 / 2}}{\mathrm{p}_{\mathrm{k}-1 / 2}} \tag{2.20}
\end{equation*}
$$

which gives

$$
\begin{equation*}
\phi_{k+1 / 2}=\phi_{\text {surf }}+\sum_{j=k+1}^{N L E V} R_{d r y}\left(T_{v}\right) \ln \frac{p_{j+1 / 2}}{p_{j-1 / 2}} \tag{2.21}
\end{equation*}
$$

where $\phi_{\text {surf }}$ is the geopotential at the surface. Full-level values of the geopotential, as required in the momentum equations (2.1) and (2.2), are given by
$\phi_{\mathrm{k}}=\phi_{\mathrm{k}+1 / 2}+\alpha_{\mathrm{k}} \mathrm{R}_{\mathrm{dry}}\left(\mathrm{T}_{\mathrm{v}}\right)_{\mathrm{k}}$
where $\alpha_{1}=\ln 2$ and, for $k>1$,

$$
\begin{equation*}
\alpha_{\mathrm{k}}=1-\frac{\mathrm{p}_{\mathrm{k}-1 / 2}}{\Delta \mathrm{p}_{\mathrm{k}}} \ln \left(\frac{\mathrm{p}_{\mathrm{k}+1 / 2}}{\mathrm{p}_{\mathrm{k}-1 / 2}}\right) \tag{2.23}
\end{equation*}
$$

The remaining part of the pressure gradient terms in (2.1) and (2.2) is given by

$$
\begin{equation*}
\mathrm{R}_{\mathrm{dry}}\left(\mathrm{~T}_{\mathrm{v}} \nabla \ln \mathrm{p}\right)_{\mathrm{k}}=\frac{\mathrm{R}_{\mathrm{dry}}\left(\mathrm{~T}_{\mathrm{v}}\right)_{\mathrm{k}}}{\Delta \mathrm{p}_{\mathrm{k}}}=\left\{\ln \left(\frac{\mathrm{p}_{\mathrm{k}+1 / 2}}{\mathrm{p}_{\mathrm{k}-1 / 2}}\right) \nabla \mathrm{p}_{\mathrm{k}-1 / 2}+\alpha_{\mathrm{k}} \nabla\left(\Delta \mathrm{p}_{\mathrm{k}}\right)\right\} \tag{2.24}
\end{equation*}
$$

with $\alpha_{k}$ given by (2.23) for all .
Finally, the energy conversion term in the thermodynamic equation (2.3) is discretized as

$$
\begin{align*}
\frac{\kappa T_{v} \omega}{(1+(\delta-1) q) p}= & \\
& \frac{\kappa\left(T_{v}\right)_{k}}{1+(\delta-1) q_{k}}\left(-\frac{1}{\Delta p_{k}}\left(\left(\ln \frac{p_{k+1 / 2}}{p_{k-1 / 2}}\right) \sum_{j=1}^{k-1}\left(D_{j} \Delta p_{j}+p_{\text {surf }}\left(\mathbf{v}_{j} \cdot \nabla \ln p_{\text {surf }}\right) \Delta B_{j j}\right)\right.\right.  \tag{2.25}\\
& \left.\left.+\alpha_{k}\left(D_{k} \Delta p_{k}+p_{\text {surf }}\left(\mathbf{v}_{k} \cdot \nabla \ln p_{\text {surf }}\right) \Delta B_{k}\right)\right\}\left(+\frac{p_{\text {surf }}}{\Delta \mathrm{p}_{\mathrm{k}}}\left\{\Delta \mathrm{~B}_{\mathrm{k}}+\frac{\mathrm{C}_{\mathrm{k}}}{\Delta \mathrm{p}_{\mathrm{k}}} \ln \frac{\mathrm{p}_{\mathrm{k}+1 / 2}}{\mathrm{p}_{\mathrm{k}-1 / 2}}\right\}\left(\mathbf{v}_{\mathrm{k}} \cdot \nabla \ln \mathrm{p}_{\text {surf }}\right)\right\}\right)
\end{align*}
$$

where $\alpha_{1}=\ln 2, \alpha_{k}$, is defined by (2.23) for $\mathrm{k}>1$, and

$$
\begin{equation*}
C_{k}=A_{k+1 / 2} B_{k-1 / 2}-A_{k-1 / 2} B_{k+1 / 2} . \tag{2.26}
\end{equation*}
$$

The reasons behind the various choices made in this vertical discretization scheme are discussed by Simmons and Burridge (1981); basically the scheme is designed to conserve angular momentum and energy, for frictionless adiabatic flow.

### 2.2.2 Finite-element vertical discretization

In CY24R3 the vertical discretization has been changed in the operational model from the finite-difference discretization in Lorenz staggering described in the previous subsection to a finite-element discretization using cubic Bsplines as basis functions.

For the finite-element (FE) discretization, all variables (even pressure) are kept at the same levels (full levels), i.e. the values of pressure at full levels and not at half levels are required. Also, the values of the derivatives $\mathrm{dA} / \mathrm{d} \eta$ and $\mathrm{dB} / \mathrm{d} \eta$ at full levels are now needed, from which the vertical derivative of pressure can be computed according to $\partial p / \partial \eta=(d A / d \eta)+(d B / d \eta) p_{s}$. In the semi-Lagrangian version of the evolution equations these are the only vertical derivatives required. They are constant in time and linked to the definition of the vertical coordinate. It is therefore convenient to change the definition of the vertical coordinate and supply $(d A / d \eta)_{k}$ and $(d B / d \eta)_{k}$ at full levels (instead of $A_{k+1 / 2}$ and $B_{k+1 / 2}$ at half levels) such that

$$
\begin{equation*}
\left(\frac{\partial \mathrm{p}}{\partial \eta}\right)_{k}=\left(\frac{\mathrm{dA}}{\mathrm{~d} \eta}\right)_{k}+\left(\frac{\mathrm{dB}}{\mathrm{~d} \eta}\right)_{k} \mathrm{p}_{\mathrm{s}} \tag{2.27}
\end{equation*}
$$

and

$$
\begin{equation*}
\int_{0}^{1} \frac{\mathrm{dA}}{\mathrm{~d} \eta} \mathrm{~d} \eta=0 \quad, \quad \int_{0}^{1} \frac{\mathrm{~dB}}{\mathrm{~d} \eta} \mathrm{~d} \eta=1 \tag{2.28}
\end{equation*}
$$

The two conditions of Eq.(2.28) above ensure that the integral of pressure from the top of the atmosphere to the surface yields exactly the surface pressure. These conditions have to be fulfilled to a good approximation with the numerical integration scheme used. Pressure at any full level can then be obtained by integrating Eq. (2.27) from the top of the atmosphere to the level in question.

The only operation in the vertical which has to be evaluated is the vertical integration. An integral operator based on a finite-element representation will be derived next.

Most of the integrals we have to evaluate are integrals from the top of the atmosphere to the individual model levels and to the surface. We therefore derive an operator in finite-element representation for this type of integral, i.e. an operator which returns the integral from the top of the atmosphere to each of the model levels $\eta_{k}$ and to the surface ( $\eta=1$ ). The vertical integral in the hydrostatic equation (i.e. from the surface upwards) can be constructed by taking the difference of the integral from the top of the atmosphere to the model level in question minus the integral from the top to the surface.
Let $\left\{d_{i}(\eta)\right\}_{i=K_{1}}^{K_{2}}$ and $\left\{e_{i}(\eta)\right\}_{i=M_{1}}^{M_{2}}$ be two sets of linearly independent functions of compact support which can be used as basis functions to expand any function of the vertical coordinate $\eta$ given in the domain $[0,1]$.

The vertical integral

$$
F(\eta)=\int_{0}^{\eta} f(x) d x
$$

can then be approximated as

$$
\begin{equation*}
\sum_{i=K_{1}}^{k_{2}} c_{i} d_{i}(\eta)=\sum_{i=M_{1}}^{M_{2}} c_{i} \int_{0}^{\eta} e_{i}(x) d x \tag{2.29}
\end{equation*}
$$

where $C_{i}$ are the coefficients of the expansion of $F(\eta)$ as a linear combination of the basis functions $\left\{d_{i}(\eta)\right\}$ and $c_{i}$ are the coefficients of the expansion of $f(\eta)$ as a linear combination of the basis functions $\left\{e_{\rho}(\eta)\right\}$.

We can then apply the Galerkin procedure to Eq. (2.29) by multiplying both sides of this equation by each function from a complete set of "test functions" $\left\{t_{j}\right\}_{j=N_{1}}^{N_{2}}$ and integrating over the vertical domain:

$$
\sum_{i=K_{1}}^{K_{2}} C_{i} \int_{0}^{1} d_{i}(x) t_{j}(x) d x=\sum_{i=M_{1}}^{M_{2}} c_{i} \int_{0}^{1}\left[t_{j}(x) \int_{0}^{x} e_{i}(y) d y\right] d x
$$

In matrix form this can be expressed as $\underset{\sim}{A} \tilde{C}=\underset{\sim}{B} \tilde{C} \Rightarrow \quad \tilde{C}={\underset{\sim}{A}}^{-1} \underset{\sim}{B} \tilde{C}$.
Incorporating into the above expression also the transformations from physical space to finite-element space and
 ical space composed mainly of the values of $f$ and $F$, respectively, at the model levels: $f_{i}=f\left(\eta_{i}\right), F_{i}=F\left(\eta_{i}\right)$, $1 \leq \mathrm{i} \leq \mathrm{N}$. The set of values $F$ also includes the value at the surface of the model. Details of the basis functions chosen to implement the scheme as well as how to compute the projection matrices $\underset{\sim}{S}$ and $\underset{\sim}{P}$ are given in Untch and Hortal (2002) .
Matrix $\underset{\sim}{I} \equiv \underset{\sim}{P}{\underset{\sim}{A}}^{-1} \underset{\sim}{B} S_{\sim}^{-1}$ is the integration operator in finite-element formulation which, applied to a function given at full model levels, yields the integrals of this function from the top of the atmosphere to each individual full model level and to the surface. All the finite sums on the vertical levels in subsection 2.2.1 are replaced by vertical integrals computed by applying the matrix integration operator $I$. Moreover the quantities $\alpha_{\mathrm{k}}$ are no longer needed as the integration operator gives directly the value of the integral at the model levels (the half levels do not have any meaning in the FE discretization).

### 2.2.3 Time discretization

To introduce a discretization in time, together with a semi-implicit correction, we define the operators

$$
\begin{aligned}
& \delta_{\mathrm{t}} \mathrm{X}=\left(\mathrm{X}^{+}-\mathrm{X}^{-}\right) / 2 \Delta \mathrm{t} \\
& \Delta_{\mathrm{tt}} \mathrm{X}=\left(\mathrm{X}^{+}-2 \mathrm{X}+\mathrm{X}^{-}\right)
\end{aligned}
$$

where $X$ represents the value of a variable at time $t, \quad X^{+}$the value at time $(t+\Delta t)$, and $X^{-}$the value at $(t-\Delta t)$. In preparation for the semi-Lagrangian treatment to be developed in section 3, we also introduce the three-dimensional advection operator

$$
\begin{equation*}
A(X)=\frac{1}{a \cos ^{2} \theta}\left(U \frac{\partial x}{\partial \lambda}+V \cos \theta \frac{\partial x}{\partial \theta}\right)+\dot{\eta} \frac{\partial x}{\partial \eta} \tag{2.30}
\end{equation*}
$$

Introducing the semi-implicit correction terms, Eqs. (2.1)-(2.4) become:

$$
\begin{align*}
& \left.\delta_{t} U+A(U)-f V+\frac{1}{a}\left\{\frac{\partial \phi}{\partial \lambda}+R_{\text {dry }} T_{v} \frac{\partial}{\partial \lambda} \ln \mathrm{p}\right)\right\}= \\
& \quad P_{U}+K_{U}-\frac{\beta}{2 a} \Delta_{\mathrm{tt}}\left\{[\gamma] \frac{\partial T}{\partial \lambda}+R_{\text {dry }} T^{\text {ref }} \frac{\partial}{\partial \lambda}\left(\ln \mathrm{p}_{\text {surf }}\right)\right\} \tag{2.31}
\end{align*}
$$

$$
\begin{gather*}
\delta_{\mathrm{t}} \mathrm{~V}+\mathrm{A}(\mathrm{~V})+\frac{\sin \theta}{\mathrm{a} \cos _{\theta}^{2}}\left(\mathrm{U}^{2}+\mathrm{V}^{2}\right)+\mathrm{fU}+\frac{\cos \theta}{\mathrm{a}}\left\{\frac{\partial \phi}{\partial \theta}+\mathrm{R}_{\mathrm{dry}} \mathrm{~T}_{\mathrm{v}} \frac{\partial}{\partial \theta}(\ln \mathrm{p})\right\}=  \tag{2.32}\\
\mathrm{P}_{\mathrm{V}}+\mathrm{K}_{\mathrm{V}}-\frac{\beta \cos \theta}{2 \mathrm{a}} \Delta_{\mathrm{tt}}\left([\gamma] \frac{\partial \mathrm{T}}{\partial \theta}+\mathrm{R}_{\mathrm{dry}} \mathrm{~T}^{\text {ref }} \frac{\partial}{\partial \theta}\left(\ln \mathrm{p}_{\text {surf }}\right)\right) \\
\delta_{\mathrm{t}} \mathrm{~T}+\mathrm{A}(\mathrm{~T})-\frac{\kappa \mathrm{T}_{\mathrm{v}} \omega}{(1+(\delta-1) \mathrm{q}) \mathrm{p}}=\mathrm{P}_{\mathrm{T}}+\mathrm{K}_{\mathrm{T}}-\frac{\beta}{2} \Delta_{\mathrm{tt}}([\tau] \mathrm{D})  \tag{2.33}\\
\delta_{\mathrm{t}} \mathrm{q}+\mathrm{A}(\mathrm{q})=\mathrm{P}_{\mathrm{q}}+\mathrm{K}_{\mathrm{q}} \tag{2.34}
\end{gather*}
$$

where $\beta$ is a parameter of the semi-implicit scheme; the classical scheme (Robert 1969) is recovered with $\beta=1$. The semi-implicit correction terms are linearized versions of the pressure gradient terms in (2.1)-(2.2) and the energy conversion term in (2.3). Thus $\mathrm{T}^{\text {ref }}$ is a reference temperature (here chosen to be independent of vertical level), while $[\gamma]$ and $[\tau]$ are matrices such that

$$
\begin{gather*}
([\gamma] \mathbf{T})_{k}=\alpha_{k}^{\text {ref }} R_{d r y} T_{k}+\sum_{j=k+1}^{N L E V} R_{d r y} T_{j} \ln \left(\frac{p_{j+1 / 2}^{\text {ref }}}{p_{j-1 / 2}^{\text {ref }}}\right),  \tag{2.35}\\
([\tau] \mathbf{D})_{k}=\kappa T^{\text {ref }}\left\{\frac{1}{\Delta p_{k}^{\text {ref }}} \ln \left(\frac{p_{k+1 / 2}^{\text {ref }}}{p_{k-1 / 2}^{\text {ref }}}\right) \sum_{j=1}^{k-1}\left(D_{j} \Delta p_{j}^{\text {ref }}\right)+\alpha_{k}^{\text {ref }} D_{k}\right\} . \tag{2.36}
\end{gather*}
$$

where the half-level pressures appearing in (2.35) and (2.35) are reference values obtained from (2.11) by choosing a reference value $p_{\text {surf }}^{\text {ref }}$ of $p_{\text {surf }}$, and the coefficients $\alpha_{k}^{\text {ref }}$ are based on these reference values. The reference values adopted for the semi-implicit scheme are $T^{\text {ref }}=300 \mathrm{~K}$ and $\mathcal{p}_{\text {surf }}^{\text {ref }}=800 \mathrm{hPa}$.

The integrated surface pressure tendency equation (2.14) becomes

$$
\begin{equation*}
\delta_{\mathrm{t}}\left(\ln \mathrm{p}_{\text {surf }}\right)+\sum_{\mathrm{k}=1}^{\mathrm{NLEV}}\left\{\frac{1}{\mathrm{p}_{\text {surf }}} \mathrm{D}_{\mathrm{k}} \Delta \mathrm{p}_{\mathrm{k}}+\left(\mathbf{v}_{\mathrm{k}} \cdot \nabla \ln \mathrm{p}_{\text {surf }}\right) \Delta \mathrm{B}_{\mathrm{k}}\right\}=-\frac{\beta}{2} \Delta_{\mathrm{tt}}[\mathrm{v}] \mathrm{D} \tag{2.37}
\end{equation*}
$$

where

$$
\begin{equation*}
[v] D=\frac{1}{p_{\text {surf }}^{\text {ref }}} \sum_{j=1}^{\text {NLEV }} D_{j} \Delta p_{\text {surf }}^{\text {ref }} \tag{2.38}
\end{equation*}
$$

### 2.2.4 Horizontal grid

A novel feature of the model is the optional use of a reduced Gaussian grid, as described by Hortal and Simmons (1991). Thus, the number of points on each latitude row is chosen so that the local east-west grid length remains approximately constant, with the restriction that the number should be suitable for the FFT ( $N=2^{p} 3^{q} 5^{r}$ ). After some experimentation, the 'fully reduced grid' option of Hortal and Simmons was implemented; all possible wavenumbers (up to the model's truncation limit) are used in the Legendre transforms. A small amount of noise in the
immediate vicinity of the poles was removed by increasing the number of grid points in the three most northerly and southerly rows of the grid (from 6, 12 and 18 points in the original design of the T213 grid to 12,16 and 20 points respectively). Courtier and Naughton (1994) have very recently reconsidered the design of reduced Gaussian grids.

### 2.2.5 Time-stepping procedure

The time-stepping procedure for the Eulerian $\mathrm{U}-\mathrm{V}$ version of the model follows closely that outlined by Temperton (1991) for the shallow-water equations. At the start of a time-step, the model state at time $(\mathrm{t}-\Delta \mathrm{t})$ is defined by the values of $\mathrm{U}, \mathrm{V}, \mathrm{T}, \mathrm{q}$ and $\ln \mathrm{p}_{\text {surf }}$ on the Gaussian grid. To compute the semi-implicit corrections, the $(t-\Delta t)$ values of divergence $D, \partial \mathrm{P} / \partial \lambda$ and $\partial \mathrm{P} / \partial \mu$ are also held on the grid, where $\mu=\sin \theta$ and

$$
\begin{equation*}
P=[\gamma] T+R_{\text {dry }} T^{\mathrm{ref}} \ln \mathrm{p}_{\text {surf }} \tag{2.39}
\end{equation*}
$$

The model state at time $t$ is defined by the spectral coefficients of $\zeta, \mathrm{D}, \mathrm{T}, \mathrm{q}$ and $\ln \mathrm{p}_{\text {surf }}$. Legendre transforms followed by Fourier transforms are then used to compute $\zeta, \mathrm{D}, \mathrm{U}, \mathrm{V}, \mathrm{T}, \partial \mathrm{T} / \partial \mu, \ln \mathrm{p}_{\text {surf }}$ and $\partial\left(\ln \mathrm{p}_{\text {surf }}\right) / \partial \mu$ at time $t$ on the model grid. Additional Fourier transforms are used to compute the corresponding values of $\partial U / \partial \lambda, \partial \mathrm{V} / \partial \lambda . \partial \mathrm{T} / \partial \lambda, \partial \mathrm{q} / \partial \lambda$ and $\partial\left(\ln \mathrm{p}_{\text {surf }}\right) / \partial \lambda$. The meridional gradients of U and V are obtained using the relationships

$$
\begin{aligned}
& \cos \theta \frac{\partial V}{\partial \theta}=a D \cos ^{2} \theta-\frac{\partial U}{\partial \lambda} \\
& \cos \theta \frac{\partial U}{\partial \theta}=\frac{\partial V}{\partial \lambda}-\zeta \cos ^{2} \theta
\end{aligned}
$$

All the information is then available to evaluate the terms at time t on the left-hand sides of (2.31)-(2.34) and (2.37), and thus to compute 'provisional' tendencies of the model variables. These tendencies (together with values of the variables at $(\mathrm{t}-\Delta \mathrm{t})$ are supplied to the physical parametrization routines, which increment the tendencies with their respective contributions. The semi-implicit correction terms evaluated at time-levels ( $\mathrm{t}-\Delta \mathrm{t}$ ) and t are then added to the tendencies. Ignoring the horizontal diffusion terms (which are handled later in spectral space), and grouping together the terms which have been computed on the grid, (2.31)-(2.34) and (2.37) can be written in the form

$$
\begin{gather*}
\mathrm{U}^{+}+\frac{\beta \Delta \mathrm{t}}{\mathrm{a}} \frac{\partial \mathrm{p}^{+}}{\partial \lambda}=\mathrm{R}_{1}  \tag{2.40}\\
\mathrm{~V}^{+}+\frac{\beta \Delta \mathrm{t}}{\mathrm{a}} \cos \theta \frac{\partial \mathrm{p}^{+}}{\partial \theta}=\mathrm{R}_{2}  \tag{2.41}\\
\mathrm{~T}^{+}+\beta \Delta \mathrm{t}[\tau] \mathrm{D}^{+}=\mathrm{R}_{3}  \tag{2.42}\\
\mathrm{q}^{+}=\mathrm{R}_{4}  \tag{2.43}\\
\left(\ln \mathrm{p}_{\text {surf }}\right)^{+}+\beta \Delta \mathrm{t}[\mathrm{v}] \mathrm{D}^{+}=\mathrm{R}_{5} \tag{2.44}
\end{gather*}
$$

The right-hand sides $\mathrm{R}_{1}-\mathrm{R}_{5}$ are transformed to spectral space via Fourier transforms followed by Gaussian integration. The curl and divergence of (2.40) and (2.41) are then computed in spectral space, leading to

$$
\begin{gather*}
\zeta^{+}=\operatorname{curl}\left(\mathrm{R}_{1}, \mathrm{R}_{2}\right)  \tag{2.45}\\
\mathrm{D}^{+}+\beta \Delta t \nabla^{2} \mathrm{p}^{+}=\operatorname{div}\left(\mathrm{R}_{1}, \mathrm{R}_{2}\right) \tag{2.46}
\end{gather*}
$$

Eqs. (2.42), (2.44) and (2.46) can then be combined with the aid of (2.39) to obtain an equation of the form

$$
\begin{equation*}
\left([I]+\frac{\mathrm{n}(\mathrm{n}+1)}{\mathrm{a}^{2}}[\Gamma]\right)\left(\mathrm{D}_{\mathrm{n}}^{\mathrm{m}}\right)^{+}=(\tilde{\mathrm{D}})_{\mathrm{n}}^{\mathrm{m}} \tag{2.47}
\end{equation*}
$$

for each zonal wavenumber m and total wavenumber n , where the matrix

$$
\begin{equation*}
[\Gamma]=\beta^{2}(\Delta \mathrm{t})^{2}\left([\gamma][\tau]+\mathrm{R}_{\mathrm{dry}} \mathrm{~T}^{\mathrm{ref}}[v]\right) \tag{2.48}
\end{equation*}
$$

couples all the NLEV values of $\left(D_{n}^{m}\right)^{+}$in a vertical column. Once $D^{+}$has been found, the calculation of $\mathrm{T}^{+}$and ( $\left.\ln \mathrm{p}_{\text {surf }}\right)^{+}$can be completed, while $\mathrm{q}^{+}$and $\zeta^{+}$have already been obtained from (2.43) and (2.45).

Finally, a 'fractional step' approach is used to implement the horizontal diffusion of vorticity, divergence, temperature and specific humidity. A simple linear diffusion of order $2 r$ is applied along the hybrid coordinate surfaces:

$$
\begin{equation*}
K_{X}=-(-1)^{r} K \nabla^{2 r} X \tag{2.49}
\end{equation*}
$$

where $X=\zeta$, $D$ or $q$. It is applied in spectral space to the $(t+\Delta t)$ values such that if $X_{n}^{m}$ is the spectral coefficient of $X$ prior to diffusion, then the diffused value $\bar{X}_{n}^{m}$ is given by

$$
\begin{equation*}
\bar{X}_{n}^{m}=\left\{1+2 \Delta t K\left(\frac{\mathrm{n}(\mathrm{n}+1)}{\mathrm{a}^{2}}\right)^{r}\right\}^{-1} X_{\mathrm{n}}^{m} \tag{2.50}
\end{equation*}
$$

A modified form of (2.50) is also used for the temperature T , to approximate diffusion on surfaces of constant pressure rather than on the sloping hybrid coordinate surfaces (Simmons, 1987). The operational version of the model uses fourth-order horizontal diffusion $(r=2)$

### 2.2.6 Time filtering

To avoid decoupling of the solutions at odd and even time steps, a Robert filter (Asselin 1972) is applied at each timestep. The time-filtering is defined by

$$
\begin{equation*}
X_{f}=X+\varepsilon\left(X_{f}^{-}-2 X+X^{+}\right) \tag{2.51}
\end{equation*}
$$

where the subscript $f$ denotes a filtered value, and $X^{-}, X$ and $X^{+}$represent values at $(t-\Delta t), t$ and $(t+\Delta t)$, respectively.

Because of the scanning structure of the model (see Chapter 4 'Computational details'), it is convenient to apply the time-filtering in grid-point space, and to split (2.51) into two parts:

$$
\begin{equation*}
\tilde{X}_{f}=X+\varepsilon\left(X_{f}^{-}-2 X\right) \tag{2.52}
\end{equation*}
$$

$$
\begin{equation*}
X_{f}=\tilde{X}_{f}+\varepsilon X^{+} \tag{2.53}
\end{equation*}
$$

The 'partially filtered' values computed by (2.52) are stored on a grid-point work file and passed from one timestep to the next. Thus, the information available after the transforms to grid-point space consists of partially filtered values at time $(\mathrm{t}-\Delta \mathrm{t})$ together with unfiltered values at time t . The filtering of the $(\mathrm{t}-\Delta \mathrm{t})$ fields can then be completed via (2.53), which after shifting by one timestep becomes:

$$
\begin{equation*}
X_{f}^{-}=\tilde{X}_{f}^{-}+\varepsilon X \tag{2.54}
\end{equation*}
$$

The computations described in Section 2.2 .5 are performed using these fully filtered values at time $(\mathrm{t}-\Delta \mathrm{t})$ and the unfiltered values at time $t$. Once (2.54) has been implemented, values of $X_{f}^{-}$are also available to implement (2.52) for the partially filtered values to be passed on to the next timestep.

### 2.2.7 Remarks

Ritchie (1988) noted that for a spectral model of the shallow-water equations, the $\mathrm{U}-\mathrm{V}$ form and the $\zeta-\mathrm{D}$ form gave identical results (apart from round-off error). In extending this work to a multi-level model, Ritchie (1991) found that this equivalence was not maintained. This was in fact a result of some analytic manipulations in the vertical, used to eliminate between the variables in solving the equations of the semi-implicit scheme, which were not exactly matched by the finite-element vertical discretization of Ritchie's model.

In the case of the model described here, the corresponding elimination between the variables is purely algebraic, and the equivalence between the $U-V$ form and the $\zeta-D$ form is maintained apart from one small exception due to the use of the hybrid vertical coordinate. In the $U-V$ model, the gradients of the geopotential $\phi$ are computed in grid-point space (from the spectrally computed gradients of $\mathrm{T}, \mathrm{q}$ and $\ln \mathrm{p}_{\text {surf }}$ ), while in the $\zeta$ - D model $\phi$ itself is computed and transformed separately into spectral space, where its Laplacian is added into the divergence equation. Since $\phi$ is not a quadratic function of the model variables there is some aliasing, which is different for the two versions of the model. In practice the differences between the $\zeta-\mathrm{D}$ model and the $\mathrm{U}-\mathrm{V}$ model were found to be very small, and in the case of a pure sigma-coordinate the two models would be algebraically equivalent.

The $\mathrm{U}-\mathrm{V}$ model is nevertheless considerably more economical than its $\zeta$ - D counterpart in terms of the number of Legendre transforms required. In addition to the transform of $\phi$ referred to above, four Legendre transforms are saved in the treatment of the wind fields using the procedures described by Temperton (1991) for the shallow-water equations. The number of multi-level Legendre transforms is thereby reduced from 17 to 12 per time-step.

### 2.2.8 $\mathrm{T}_{\mathrm{v}}$ as spectral variable

In preparation for a further reduction in the number of Legendre transforms required by the semi-Lagrangian version of the model, the modified Eulerian version includes an option to keep the virtual temperature $\mathrm{T}_{\mathrm{v}}$, rather than the temperature T , as the spectral variable. In the time-stepping procedure, Legendre transforms followed by Fourier transforms are used to compute $T_{v}, \partial T_{v} / \partial \mu$ and $\partial T_{v} / \partial \lambda$ at time $t$ on the model grid; the corresponding values of $T, \partial T / \partial \mu$ and $\partial T / \partial \lambda$ are then computed using the corresponding values of $q, \partial q / \partial \mu$ and $\partial q / \partial \lambda$. The thermodynamic equation (2.3) is then stepped forward in time exactly as before. After the physical parametrization routines, the 'provisional' value of $\mathrm{T}(\mathrm{t}+\Delta \mathrm{t})$ is combined with $\mathrm{q}(\mathrm{t}+\Delta \mathrm{t})$ to compute a provisional value of $\mathrm{T}_{\mathrm{v}}(\mathrm{t}+\Delta \mathrm{t})$. The semi-implicit correction terms evaluated at time-levels $(\mathrm{t}-\Delta \mathrm{t})$ and $\mathrm{T}_{\mathrm{v}}(\mathrm{t}+\Delta \mathrm{t})$ are then added to the provisional value of $\mathrm{T}_{\mathrm{v}}(\mathrm{t}+\Delta \mathrm{t})$, just before the transform back to spectral space.

There are corresponding slight changes in the semi-implicit correction terms. The linearized hydrostatic matrix [ $\gamma$ ]
in (2.31)-(2.32) and (2.39) now operates on $\mathrm{T}_{\mathrm{v}}$ rather than on T . From the point of view of the semi-implicit scheme, (2.33) has implicitly been replaced by an equation of the form

$$
\begin{equation*}
\delta_{\mathrm{t}} \mathrm{~T}_{\mathrm{v}}=\ldots-\frac{\beta}{2} \Delta_{\mathrm{tt}}([\tau] \mathrm{D}) \tag{2.55}
\end{equation*}
$$

although as explained above it is not necessary to formulate or compute the missing terms explicitly. Hence, (2.42) is replaced by

$$
\begin{equation*}
\mathrm{T}_{\mathrm{v}}^{+}+\beta \Delta \mathrm{t}[\tau] \mathrm{D}^{+}=\mathrm{R}_{3}^{\prime} \tag{2.56}
\end{equation*}
$$

and the solution of the semi-implicit equations in spectral space proceeds just as before.
This change of spectral variable results in only insignificant changes to a 10-day model forecast, but permits useful economies in the semi-Lagrangian version to be described in the next chapter.

## IFS documentation

## Part III: Dynamics and numerical procedures

## CHAPTER 3 Semi-Lagrangian formulation

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### 3.1 GENERAL DESCRIPTION

The general form of the model equations is

$$
\begin{equation*}
\frac{\mathrm{d} X}{\mathrm{dt}}=\frac{\partial \mathrm{X}}{\partial \mathrm{t}}+\mathrm{A}(\mathrm{X})=\mathrm{R}=\mathrm{L}+\mathrm{N} \tag{3.1}
\end{equation*}
$$

where the three-dimensional advection operator A was defined in (2.30), L is the linearized part of R and N is the remainder or "non-linear terms". An explicit three-time-level semi-Lagrangian treatment of (3.1) is obtained by finding the approximate trajectory, over the time interval $[t-\Delta t, t+\Delta t]$, of a particle which arrives at each grid point $\underset{\sim}{\mathrm{X}}$ at time $(\mathrm{t}+\Delta \mathrm{t})$. Equation (3.1) is then approximated by

$$
\begin{equation*}
\frac{\mathrm{X}^{+}-\mathrm{X}^{-}}{2 \Delta \mathrm{t}}=\mathrm{R}^{0} \tag{3.2}
\end{equation*}
$$

where the superscripts,+ 0 and - , respectively denote evaluation at the arrival point $(\underset{\sim}{x}, \mathrm{t}+\Delta \mathrm{t})$, the mid-point
of the trajectory $(\underset{\sim}{\mathrm{X}}-\underset{\sim}{\alpha}, \mathrm{t})$, and the departure point $(\underset{\sim}{\mathrm{X}}-\underset{\sim}{\alpha}, \mathrm{t}-\Delta \mathrm{t})$. Since the mid-point and the departure point will not in general coincide with model grid points, $X^{-}$and $R^{0}$ must be determined by interpolation.

It is more economical (and, as discussed later, gives better results in some circumstances; see also Tanguay et al., 1992) to evaluate the right-hand side of (3.2) as

$$
\begin{equation*}
\mathrm{R}^{0}=\frac{1}{2}\{\mathrm{R}(\underset{\sim}{\mathrm{x}}-2 \underset{\sim}{\alpha}, \mathrm{t})+\mathrm{R}(\underset{\sim}{\mathrm{x}}, \mathrm{t})\} \tag{3.3}
\end{equation*}
$$

since only a single interpolation (of the combined field $X(t-\Delta t)+\Delta t R(t)$ at the point $(\underset{\sim}{x}-\underset{\sim}{\alpha})$ ) is then required in order to determine $X^{+}$.

The right-hand sides of the time-discretized model equations also contain semi-implicit correction terms, which in the Eulerian model took the form

$$
\Delta_{\mathrm{tt}} \mathrm{X}=\left(\mathrm{X}^{+}-2 \mathrm{X}^{0}+\mathrm{X}^{-}\right)
$$

where the superscripts refer to time-levels, and to a single common grid point. In the semi-Lagrangian version of the model, the semi-implicit correction terms take the form

$$
\begin{equation*}
\Delta_{\mathrm{tt}} \mathrm{X}=(\mathrm{X}(\underset{\sim}{\mathrm{x}}, \mathrm{t}+\Delta \mathrm{t})-\mathrm{X}(\underset{\sim}{\mathrm{x}}, \mathrm{t}))+(\mathrm{X}(\underset{\sim}{\mathrm{X}}-2 \underset{\sim}{\alpha}, \mathrm{t}-\Delta \mathrm{t})-\mathrm{X}(\underset{\sim}{\mathrm{X}}-2 \underset{\sim}{\alpha}, \mathrm{t})) \tag{3.4}
\end{equation*}
$$

and again the terms to be evaluated at the departure point $(\underset{\sim}{\mathrm{X}}-\underset{\sim}{\alpha})$ can be added to other right-hand side terms before interpolation. Notice that the evaluation of $\Delta_{t t} X$, and both ways of evaluating $R^{0}$, are all centred in space and time.

To obtain accurate results from a semi-Lagrangian integration scheme, it is necessary to choose the order of interpolation carefully (see for example Staniforth and Côté, 1991). In practice it has been found (for the model described here) that linear interpolation is adequate for the terms evaluated at the midpoint of the trajectory, but that cubic interpolation is essential for the terms evaluated at the departure point. Cubic interpolation in three dimensions is expensive, and fortunately a 'quasi-cubic' interpolation (suggested by Courtier) was found to give essentially equivalent results. The technique can be illustrated by two-dimensional interpolation on a regular grid. The target point is at $\left(x_{1}+\alpha, y_{j}+\beta\right)$. In the first step, four interpolations are performed in the $x$-direction: linear (rather than the usual cubic) interpolations to the points $\left(\mathrm{X}_{\mathrm{I}}+\alpha, \mathrm{y}_{\mathrm{J}-1}\right)$ and $\left(\mathrm{X}_{\mathrm{I}}+\alpha, \mathrm{y}_{\mathrm{J}+2}\right)$, and cubic interpolations to the points $\left(x_{I}+\alpha, y_{J}\right)$ and $\left(x_{I}+\alpha, y_{J+1}\right)$. In the second step, one cubic interpolation is performed in the $y$-direction, to evaluate the field at the target point. The number of 'neighbours' contributing to the result is reduced from 16 to 12 . The generalization to three dimensions is straightforward and results in a significant saving, the number of neighbours being reduced from 64 to 32 , and the computation being reduced from 21 one-dimensional cubic interpolations to 7 cubic plus 10 linear one-dimensional interpolations.

For the reduced Gaussian grid described in Subsection 2.2.4, the mesh is no longer regular. However, it is easily seen that the extra complication is relatively minor provided that the first step in the interpolation is performed in the $\lambda$-direction.

The order of the interpolation in the vertical is reduced to linear when the evaluation point lies between the two highest model levels, or between the lowest two model levels. Extrapolation beyond the top or bottom levels is not allowed.

### 3.2 Finding the departure point

Extending the procedure of Robert (1981) to three dimensions, the midpoint $(\underset{\sim}{X}-\underset{\sim}{\alpha})$ and the departure point $(\underset{\sim}{\mathrm{X}}-2 \underset{\sim}{\alpha})$ of the trajectory for each arrival point $\underset{\sim}{\mathrm{X}}$ are found by iteratively solving the equation

$$
\begin{equation*}
\underset{\sim}{\alpha}=\Delta \operatorname{tr} \underset{\sim}{\mathbf{v}}(\underset{\sim}{\mathrm{X}}-\underset{\sim}{\alpha}, \mathrm{t}) \tag{3.5}
\end{equation*}
$$

where $\underset{\sim}{\mathbf{v}}$ in (3.5) is the three-dimensional wind field $(\mathbf{u}, \mathbf{v}, \dot{\eta})$. Since $\dot{\eta}$ was never explicitly required in the Eulerian version of the model (see Eqs. (2.18)-(2.19) for the Eulerian discretization of vertical advection), it is necessary to construct this field for the trajectory calculations. As $\dot{\eta}$ is already specified at the upper and lower boundaries $(\dot{\eta}=0$, at $\eta=0$ and at $\eta=1$ ) it would be natural to construct $\dot{\eta}$ at the half-levels (i.e. vertically staggered with respect to $u$ and $v$ ), and indeed a preliminary version of the model was coded that way. However, it is more convenient to hold the three velocity components at the same set of points (which also coincide with the arrival points), so the formulation was changed to use $\dot{\eta}$ at the 'full' levels. Thus, the vertical velocity used in (3.5) is defined by

$$
\begin{equation*}
\dot{\eta}_{k}=\frac{\frac{1}{2}\left[\left(\dot{\eta} \frac{\partial p}{\partial \eta}\right)_{k-\frac{1}{2}}+\left(\dot{\eta} \frac{\partial p}{\partial \eta}\right)_{k+\frac{1}{2}}\right]}{\left(\frac{\partial p}{\partial \eta}\right)_{k}} \tag{3.6}
\end{equation*}
$$

where $\dot{\eta} \partial \mathrm{p} / \partial \eta$ is already defined by (2.18) and

$$
\begin{equation*}
\left(\frac{\partial p}{\partial \eta}\right)_{k}=\frac{\Delta p_{k}}{\Delta \eta_{k}}=p_{\text {surf }} \frac{\Delta A_{k} / p_{\text {surf }}+\Delta B_{k}}{\Delta A_{k} / p_{0}+\Delta B_{k}} \tag{3.7}
\end{equation*}
$$

In deriving (3.7) we have used (2.11) together with a formal definition of $\eta$ itself (which again was not required by the discretized Eulerian dynamics):

$$
\begin{equation*}
\eta_{\mathrm{k}+}=\mathrm{A}_{\mathrm{k}+} / \mathrm{p}_{0}+\mathrm{B}_{\mathrm{k}+} \tag{3.8}
\end{equation*}
$$

where $\mathrm{p}_{0}$ is a constant pressure (chosen to be 1013.25 hPa ).
The iterative procedure for solving (3.5) is analogous to that used by Ritchie (1991) in a $\sigma$-coordinate model. Given an estimate ${\underset{\sim}{\alpha}}^{(\mathrm{k})}$ after k iterations, the next iteration is given by

$$
\begin{equation*}
{\underset{\sim}{\alpha}}^{(\mathrm{k}+1)}=\Delta \underset{\sim}{\operatorname{tr}}\left(\underset{\sim}{\mathrm{X}}-{\underset{\sim}{\alpha}}^{(\mathrm{k})}, \mathrm{t}\right) \tag{3.9}
\end{equation*}
$$

where the vertical $(\eta)$ component of the displacement is found first. The vertical component of ${\underset{\sim}{\alpha}}^{(k)}$ on the righthand side of (3.9) is then updated before the horizontal components are found taking into account the spherical geometry following Ritchie (1987, 1988). The first guess is given by

$$
\begin{equation*}
{\underset{\sim}{\alpha}}^{(0)}=\Delta \underset{\sim}{t}(\underset{\sim}{x}, \mathrm{t}) \tag{3.10}
\end{equation*}
$$

The calculations include approximations to the spherical geometry away from the poles, following Ritchie and Beaudoin (1994). In agreement with previous work (reviewed by Staniforth and Côté, 1991), little sensitivity was found to the order of interpolation used in the trajectory calculations, and linear interpolation appears to be sufficiently accurate. After providing a first guess via (3.10), a single further iteration was found to be adequate.

Once the midpoint $(\underset{\sim}{X}-\underset{\sim}{\alpha})$ of the trajectory has been found, the departure point $(\underset{\sim}{X}-\underset{\sim}{\alpha})$ is immediately obtained (in the horizontal, the backward extension of the trajectory is along a great circle). In the vertical, if the departure point is then above the first (or below the last) mode level, it is modified to lie on the first (last) level.

In solving (3.9), it is necessary to convert between a displacement in terms of the spatial coordinates and the corresponding displacement in terms of 'grid lengths', in order to select the correct three-dimensional block of points for the interpolation routine. This is simple in the horizontal, since the mesh length is constant in the $\lambda$-direction (at a given latitude), and almost constant in the $\theta$-direction. It is more difficult in the vertical, where the grid spacing changes rapidly, and the conversion algorithm for the vertical displacement makes use of an auxiliary grid defined with high uniform resolution.

## 3.3 'NON-INTERPOLATING’ SCHEME IN THE VERTICAL

An alternative formulation of the semi-Lagrangian scheme in three dimensions was suggested by Ritchie (1991). Equation (3.1) can be rewritten as

$$
\begin{equation*}
\frac{d_{H} X}{d t}+\dot{\eta} \frac{\partial X}{\partial \eta}=R-\dot{\eta} \frac{\partial X}{\partial \eta}+\dot{\eta} \frac{\partial x}{\partial \eta} \tag{3.11}
\end{equation*}
$$

where

$$
\frac{\mathrm{d}_{\mathrm{H}} \mathrm{X}}{\mathrm{dt}}=\frac{\partial \mathrm{X}}{\partial \mathrm{t}}+\mathrm{A}_{\mathrm{H}}(\mathrm{X})
$$

and $A_{H}$ is the horizontal part of the advection operator defined in (2.30). In (3.11), $\dot{\eta}^{*}$ is defined to be a vertical velocity which would lead to the departure point of the trajectory at time ( $t-\Delta t$ ) lying exactly on a model level. This model level is chosen to be the one closest to the true departure point. Equation (3.11) is then approximated by

$$
\begin{equation*}
\frac{X^{+}-X^{-}}{2 \Delta t}=\left(R-\dot{\eta} \frac{\partial X}{\partial \eta}\right)^{0}+\dot{\eta}^{*}\left(\frac{\partial X}{\partial \eta}\right)^{0} \tag{3.12}
\end{equation*}
$$

where the superscripts,+ 0 , - respectively denote evaluation at the arrival point $(\underset{\sim}{x}, \mathrm{t}+\Delta \mathrm{t})$, the midpoint $(\underset{\sim}{\mathrm{X}}-\underset{\sim}{\alpha}, \mathrm{t})$ and the departure point $(\underset{\sim}{\mathrm{X}}-\underset{\sim}{\alpha}, \mathrm{t}-\Delta \mathrm{t})$ of the modified trajectory. Since the modified departure point lies by definition on a model level, no vertical interpolation is required to evaluate $X^{-}$. As discussed in Subsection 3.1 above, it is also possible to evaluate the terms on the right-hand side of (3.12) by averaging the values at $(\underset{\sim}{\mathrm{X}}-2 \underset{\sim}{\alpha}, \mathrm{t})$ and $(\underset{\sim}{\mathrm{X}}, \mathrm{t})$; in this case no vertical interpolation at all is required. Notice that a separate interpolation is required to evaluate the second term on the right-hand side of (3.12) since the quantity $\dot{\eta}^{*}$, defined by

$$
\begin{equation*}
\dot{\eta}^{*}=\frac{\eta^{+}-\eta^{-}}{2 \Delta t} \tag{3.13}
\end{equation*}
$$

where $\eta^{+}$and $\eta^{-}$are respectively the arrival and departure levels of the modified trajectory, is meaningful only at each grid point.

If the vertical velocity (or the time-step) is sufficiently small, then the modified departure point lies on the same model level as the arrival point, $\dot{\eta}^{*}$ is zero and the treatment of vertical advection becomes purely Eulerian. In general there is an Eulerian treatment of the advection by the 'residual vertical velocity' $(\dot{\eta}-\dot{\eta}$ "), which is small enough to guarantee that the Eulerian CFL criterion for vertical advection is respected. Thus, the 'non-interpolat-
ing' scheme maintains the desirable stability properties of the 'fully interpolating' scheme.
There is a subtle, but important, difference in the way the iterative scheme (3.9) is implemented to determine the modified trajectory in the non-interpolating scheme. As before, the first step at each iteration is to update the estimate of the vertical component of the displacement.The implied updated departure point is then moved to the closest model level. In the second step, the horizontal components are then updated using the winds evaluated at the midpoint of the modified trajectory. Notice that this gives a result different from that obtained by simply carrying out the trajectory calculation of the fully interpolating scheme and then projecting the departure point to the nearest model level. The modified procedure described above is easily seen to be more consistent by considering the case in which the vertical velocity is not zero, but is small enough for the modified trajectory to be horizontal $\left(\dot{\eta}^{*}=0\right)$ ). The discretization is then equivalent to a purely two-dimensional semi-Lagrangian scheme, the trajectory being computed using the horizontal wind field evaluated on a single model level.

An incidental advantage of the 'non-interpolating' scheme over the 'fully interpolating' scheme is that it resolves any ambiguities about the treatment of departure points above the top model level or below the bottom model level; the modified departure points automatically lie on the top or bottom level. The treatment of vertical advection becomes Eulerian, which is well-defined at the top and bottom levels. Thus, the non-interpolating scheme removes the need for artificial 'nudging' of the departure points or the extrapolation of quantities to points above or below the domain of the model levels.

Smolarkiewicz and Rasch (1991) have extended the principle of the 'non-interpolating' semi-Lagrangian formulation to generate a broader class of stable and accurate advection schemes.

### 3.4 SEMI-LAGRANGIAN DISCRETIZATION

Here we describe in detail only the fully interpolating version of the semi-Lagrangian discretization; the modifications necessary for the 'non-interpolating in the vertical' version become evident by comparing the right-hand side of (3.12) with that of (3.2).

Following Ritchie (1988, 1991), the momentum equations are integrated in vector form to avoid an instability of the metric term near the poles. Using the notation of (3.2) and defining the horizontal wind vector $\mathbf{v}_{\mathrm{H}}=(\mathrm{u}, \mathrm{v})$, the semi-Lagrangian equivalent of (2.31)-(2.32) is

$$
\begin{equation*}
\frac{\mathbf{v}_{H}^{+}-\mathbf{v}_{\mathrm{H}}^{-}}{2 \Delta \mathrm{t}}+\left[\mathrm{fk} \times \mathbf{v}_{\mathrm{H}}+\nabla \phi+\mathrm{R}_{\mathrm{dry}} \mathrm{~T}_{\mathrm{v}} \nabla \ln \mathrm{p}\right]^{0}=\mathrm{P}_{\mathrm{v}}+\mathrm{K}_{\mathrm{v}}-\frac{\beta}{2} \Delta_{\mathrm{tt}} \nabla\left\{[\gamma] \mathrm{T}+\mathrm{R}_{\mathrm{dry}} \mathrm{~T}_{\mathrm{v}} \ln \mathrm{p}_{\text {surf }}\right\} \tag{3.14}
\end{equation*}
$$

where $\mathbf{k}$ is the vertically directed unit vector and $\nabla$ is the horizontal gradient operator in spherical coordinates. On the right-hand side of (3.14), $\mathrm{P}_{\mathrm{v}}$ and $\mathrm{K}_{\mathrm{v}}$ respectively denote the contributions of the physical parametrization schemes and horizontal diffusion, to be discussed in Subsection 3.6, while the semi-implicit correction terms are evaluated as in (3.4). For the momentum equations, it was found advantageous to evaluate the time-level $t$ terms [ ] ${ }^{0}$ as an average between the values at the departure and arrival points of the trajectory, as in (3.3). The pressure gradient terms are discretized in exactly the same way as for the Eulerian model (see Subsection 2.2.1).

Since (3.14) is in vector form, it is important to account for the change in the orientation of the coordinate system as the particle follows the trajectory; the manipulations required are as set out by Ritchie (1988) and simplified by Ritchie and Beaudoin (1994).

The thermodynamic and moisture equations (2.33)-(2.34) become

$$
\begin{gather*}
\frac{\mathrm{T}^{+}-\mathrm{T}^{-}}{2 \Delta \mathrm{t}}+-\left\{\frac{\kappa \mathrm{T}_{\mathrm{v}} \omega}{(1+(\delta-1) \mathrm{q}) \mathrm{p}}\right\}^{0}=\mathrm{P}_{\mathrm{T}}+\mathrm{K}_{\mathrm{T}}-\frac{\beta}{2} \Delta_{\mathrm{tt}}[[\tau] \mathrm{D})  \tag{3.15}\\
\frac{\mathrm{q}^{+}-\mathrm{q}^{-}}{2 \Delta \mathrm{t}}=\mathrm{P}_{\mathrm{q}}+\mathrm{K}_{\mathrm{q}} \tag{3.16}
\end{gather*}
$$

In (3.15), the $\left\}^{0}\right.$ term is discretized as in (2.25), and evaluated at the midpoint of the trajectory, while the semiimplicit correction terms are evaluated as in (3.4).

The $\eta$-coordinate continuity equation (2.5) can be rewritten as

$$
\begin{equation*}
\frac{d}{d t}\left(\frac{\partial p}{\partial \eta}\right)+\frac{\partial p}{\partial \eta}\left(D+\frac{\partial \dot{\eta}}{\partial \eta}\right)=0 \tag{3.17}
\end{equation*}
$$

Setting

$$
p=A(\eta)+B(\eta) p_{\text {surf }}
$$

and noting that

$$
\frac{\partial}{\partial t}\left(\frac{\partial A}{\partial \eta}\right)=\nabla \cdot\left(\frac{\partial A}{\partial \eta}\right)=\frac{\partial p_{\text {surf }}}{\partial \eta}=0
$$

we also have

$$
\begin{equation*}
\frac{d}{d t}\left(\frac{\partial p}{\partial \eta}\right)=\frac{\partial B}{\partial \eta} \frac{d p_{\text {surf }}}{d t}+\dot{\eta} \frac{\partial}{\partial \eta}\left(\frac{\partial p}{\partial \eta}\right) \tag{3.18}
\end{equation*}
$$

Combining (3.17) and (3.18).

$$
\begin{equation*}
\frac{\partial B}{\partial \eta} \frac{d p_{\text {surf }}}{d t}+\frac{\partial p}{\partial \eta} D+\frac{\partial}{\partial \eta}\left(\dot{\eta} \frac{\partial p}{\partial \eta}\right)=0 \tag{3.19}
\end{equation*}
$$

Now introducing the vertical discretization, (3.19) becomes

$$
\begin{equation*}
\Delta B_{k} \frac{d p_{\text {surf }}}{d t}+\Delta p_{k} D_{k}+\left(\dot{\eta} \frac{\partial p}{\partial \eta}\right)_{k+\frac{1}{2}}-\left(\dot{\eta} \frac{\partial p}{\partial \eta}\right)_{k-\frac{1}{2}}=0 \tag{3.20}
\end{equation*}
$$

the vertical discretization of $\dot{\eta} \partial \mathrm{p} / \partial \eta$ having been defined in (2.18).
Changing the prognostic variable to $\dot{\eta} \partial \rho / \partial \eta$,

$$
\begin{equation*}
\Delta B_{k} \frac{d}{d t}\left(\ln p_{\text {surf }}\right)+\frac{1}{p_{\text {surf }}}\left\{\Delta p_{k} D_{k}+\left(\dot{\eta} \frac{\partial p}{\partial \eta}\right)_{k+\frac{1}{2}}-\left(\dot{\eta} \frac{\partial p}{\partial \eta}\right)_{k+\frac{1}{2}}\right\}=0 \tag{3.21}
\end{equation*}
$$

Combining (3.21) with the discrete definition of $\dot{\eta} \partial \mathrm{p} / \partial \eta$ given by (2.18),

$$
\begin{equation*}
\Delta \mathrm{B}_{\mathrm{k}} \frac{\mathrm{~d}}{\mathrm{dt}}\left(\ln \mathrm{p}_{\text {surf }}\right)-\Delta \mathrm{B}_{\mathrm{k}}\left\{\frac{\partial\left(\ln \mathrm{p}_{\text {surf }}\right)}{\partial \mathrm{t}}+\mathbf{v}_{\mathrm{k}} \cdot \nabla \ln \mathrm{p}_{\text {surf }}\right\}=0 \tag{3.22}
\end{equation*}
$$

where $\partial\left(\ln \mathrm{p}_{\text {surf }}\right) / \partial \mathrm{t}$ is given by $(2.14)$.
Noting that

$$
\sum_{\mathrm{k}=1}^{\mathrm{NLEV}} \Delta \mathrm{~B}_{\mathrm{k}}=1
$$

and including the semi-implicit correction terms, the semi-Lagrangian discretization of the continuity equation finally becomes

$$
\begin{equation*}
\left(\ln \mathrm{p}_{\text {surf }}\right)^{+}=\sum_{\mathrm{k}=1}^{\mathrm{NLEV}} \Delta \mathrm{~B}_{\mathrm{k}}\left[\left(\ln \mathrm{p}_{\text {surf }}\right)^{-}+2 \Delta \mathrm{t}\left\{\frac{\partial\left(\ln \mathrm{p}_{\text {surf }}\right)}{\partial \mathrm{t}}+\mathbf{v}_{\mathrm{k}} \cdot \nabla \ln \mathrm{p}_{\text {surf }}\right\}^{0}-\frac{\beta \Delta \mathrm{t}}{\mathrm{p}_{\text {surf }}^{\text {ref }}} \Delta_{\mathrm{tt}}\left\{\sum_{\mathrm{j}=1}^{\mathrm{NLEV}}\left(\Delta \mathrm{p}_{\mathrm{j}}^{\text {ref }} \mathrm{D}_{\mathrm{j}}\right)\right\}\right] \tag{3.23}
\end{equation*}
$$

(Since there is no vertical advection term in (3.23), no modification is required for the vertically non-interpolating scheme). It is important to bear in mind that each contribution to the sum on the right-hand side of (3.23) involves a different trajectory. The interpolations for $\left(\ln \mathrm{p}_{\text {surf }}\right)^{-}$and the semi-implicit correction terms are however twodimensional, since these quantities are independent of vertical level. The $\left\}^{0}\right.$ term is evaluated at the midpoint of the trajectory, and requires a three-dimensional interpolation.

In summary, the semi-Lagrangian discretization is given by Equations (3.14)-(3.16)together with (3.23).

### 3.5 COMPARISON WITH OTHER SCHEMES

The semi-Lagrangian formulation presented above differs in some respects from those proposed by other authors. Perhaps the most notable difference lies in the treatment of the conversion $\omega$ term in the thermodynamic equation (3.15), and of the right-hand side of the continuity equation (3.23). Both involve terms of the form $\mathbf{v} . \nabla \ln \mathrm{p}_{\text {surf }}$, which in our scheme are computed in a purely Eulerian fashion. This may appear somewhat inconsistent; indeed McDonald and Haugen (1993) state as a specific design objective of their scheme that the operator $\mathbf{v} . \nabla$ should not appear explicitly. The alternative approach, also taken by Williamson and Olson (1994), is to use the continuity equation in its semi-implicit semi-Lagrangian form to derive a consistent equation for predicting $\dot{\eta} \partial \mathrm{p} / \partial \eta$, which can then be used to eliminate the $\mathbf{v} . \nabla \ln \mathrm{p}_{\text {surf }}$ terms. In the $\sigma$-coordinate system, Bates et al. (1993) and McDonald and Haugen (1992) used a similar approach to derive a prognostic equation for $\sigma$. A possible disadvantage of such an approach is that $\dot{\eta} \partial \mathrm{p} / \partial \eta$ (or $\sigma$ ) then follows an independent evolution, no longer satisfying a diagnostic relationship of the form (2.18). Our 'Eulerian' treatment of the $\mathbf{v} . \nabla \ln \mathrm{p}_{\text {surf }}$ terms avoids this disadvantage and seems to work well, but further study is required to determine whether this difference in formulation is important or not.

Another aspect of our semi-Lagrangian discretization of the continuity equation, which differs from that in other models, concerns the definition of the trajectory; in our scheme this is the same (three-dimensional) trajectory as used for the other variables. In the continuous form of the equation, (3.19), the advective part of the total derivative $d p / d t$ may be regarded either as two-dimensional or as three-dimensional (since $\partial p_{\text {surf }} / \partial \eta$ is zero). However the vertically discretized form, (3.20), is well-defined only at discrete model levels, implying that for consistency the semi-Lagrangian discretization (3.23) should be based on horizontal trajectories. Correcting this inconsistency
in our scheme by computing horizontal trajectories for the continuity equation, based on the horizontal wind at each model level, made very little difference to the results, and for the time being we have allowed the inconsistency to remain. (As discussed later, in the case of the 'vertically non-interpolating' scheme the modified trajectories are nearly always horizontal anyway.) In the case of the fully interpolating scheme, recomputing the trajectories represents a significant expense; Bates et al. (1993) and McDonald and Haugen (1992) used a simple projection of the three-dimensional trajectory onto the model level of the arrival point. In our model this approach resulted in poor mass conservation, though Bates et al. (1993) came to the opposite conclusion. Again, the importance or otherwise of these differences in formulation is not yet firmly established.

### 3.6 TiME-STEPPING PROCEDURE

The general outline of the time-stepping procedure for the semi-Lagrangian version is similar to that described for the Eulerian model in Subsection 2.2.5. Thus at the start of a timestep, the model state at time $(\mathrm{t}-\Delta \mathrm{t})$ is defined by the values of $U, V, T, q$ and $\ln p_{\text {surf }}$ on the Gaussian grid. To complete the semi-implicit corrections, the $(t-\Delta t)$ values of $D, \partial p / \partial \lambda$ and $\partial p / \partial \mu$ are also held on the grid. The model state at time $t$ is defined by the spectral coefficients of $\zeta, \mathrm{D}, \mathrm{T}, \mathrm{q}$ and $\ln \mathrm{p}_{\text {surf }}$. Legendre transforms followed by Fourier transforms are then used to compute $\mathrm{D}, \mathrm{U}, \mathrm{V}, \mathrm{T}, \partial \mathrm{T} / \partial \mu, \mathrm{q}, \partial \mathrm{q} / \partial \mu, \ln \mathrm{p}_{\text {surf }}$, and $\partial\left(\ln \mathrm{p}_{\text {surf }}\right) / \partial \mu$ at time t on the model grid; additional Fourier transforms are used to compute the corresponding values of $\partial \mathrm{T} / \partial \lambda, \partial q / \partial \lambda$, and $\partial\left(\ln \mathrm{p}_{\text {surf }}\right) / \partial \lambda$. Since $\zeta$ and the horizontal gradients of U and V are no longer required on the model grid, one multi-level Legendre transform and three multi-level Fourier transforms are saved in comparison with the Eulerian version.

Since the advection of moisture is handled by the semi-Lagrangian discretization (3.16), the horizontal gradients of $q$ are only needed in order to compute the horizontal gradients of the virtual temperature $T_{v}$ (which in turn are required to compute the $\nabla \phi$ term in (3.14)). If $\mathrm{T}_{\mathrm{v}}$ is chosen as the spectral variable as in Subsection 2.2.8, these gradients are available directly, and there is then no need to transform $\partial q / \partial \mu$ (or $\partial q / \partial \lambda$ ) to the model grid. The number of multi-level Legendre transforms per time-step is further reduced to 10. In passing, all the ingredients are then in place for a semi-Lagrangian treatment in which the moisture field is never transformed to spectral space (Williamson and Rasch, 1994), and only 8 multi-level Legendre transforms are required per time-step (compared with 17 in the original $\zeta$-D Eulerian model).

After the transforms to the model grid, all the information is then available to compute the trajectories for each grid point, and to evaluate the 'dynamical' contributions to the semi-Lagrangian discretization. Ignoring for a moment the contributions of the physical parametrization schemes and of the horizontal diffusion, each equation is either of the form

$$
\begin{equation*}
X^{+}(\underset{\sim}{x})=X^{-}(\underset{\sim}{X}-2 \underset{\sim}{\alpha})+\Delta t\left\{R^{0}(\underset{\sim}{X}-2 \underset{\sim}{\alpha})+R^{0}(\underset{\sim}{X})\right\}+S^{-}(\underset{\sim}{X}-2 \underset{\sim}{\alpha})+S^{+}(\underset{\sim}{X}) . \tag{3.24}
\end{equation*}
$$

or

$$
\begin{equation*}
X^{+}(\underset{\sim}{x})=X^{-}(\underset{\sim}{x}-2 \underset{\sim}{\alpha})+2 \Delta \mathrm{tR}^{0}+S^{-}(\underset{\sim}{x}-2 \underset{\sim}{\alpha})+S^{+}(\underset{\sim}{x}) \tag{3.25}
\end{equation*}
$$

depending on whether the $R^{0}$ terms are averaged between the end points of the trajectory or evaluated at the midpoints. In (3.24) and (3.25), the $S$ terms represent the semi-implicit corrections; $\mathrm{S}^{-}$includes contributions from time-levels $(\mathrm{t}-\Delta \mathrm{t})$ and t , while $\mathrm{S}^{+}$includes contributions from time-levels t and $(\mathrm{t}+\Delta \mathrm{t})$.

In the first part of the calculation for equations of the form (3.24), the combined field $\mathrm{X}^{+}+\Delta \mathrm{tR}{ }^{0}+\mathrm{S}^{-}$is computed, and the value of this combined field at each departure point $(\underset{\sim}{X}-\underset{\sim}{\alpha})$ is then found by interpolation. Adding
the (uninterpolated) value of $\Delta t R^{0}$ results in a provisional value of $X^{+}$at each grid point, incorporating all the terms in (3.24) except for $S^{+}$. The calculation for equations of the form (3.25) proceeds similarly, except that two interpolations are required, one for $\mathrm{X}^{-}+\mathrm{S}^{-}$at $(\underset{\sim}{\mathrm{X}}-2 \underset{\sim}{\alpha})$, and one for $2 \Delta \mathrm{t} \mathrm{R}^{0}$ at $(\underset{\sim}{\mathrm{X}} \underset{\sim}{\alpha})$

A provisional value $\mathrm{X}^{+}$is now available at each grid point for each variable, and is used together with $\mathrm{X}^{-}$at the same grid point to compute an 'Eulerian' tendency. These fields and their tendencies are then supplied to the physical parametrization routines, which increment the tendencies with their respective contributions, just as in the Eulerian version (except that, to avoid extra interpolations, the $\mathrm{S}^{-}$terms have been included in the supplied dynamical tendencies). If $\mathrm{T}_{\mathrm{v}}$ is chosen as the spectral variable, a provisional value of $\mathrm{T}_{\mathrm{v}}{ }^{+}$is computed at this point.

The contributions from the $S^{-}$terms at time $t$ are now added in, resulting in a set of equations of the form

$$
\begin{gather*}
\mathrm{U}^{+}+\frac{\beta \Delta \mathrm{t}}{\mathrm{a}} \frac{\partial \mathrm{p}^{+}}{\partial \lambda}=\mathrm{Q}_{1}  \tag{3.26}\\
\mathrm{~V}^{+}+\frac{\beta \Delta \mathrm{t}}{\mathrm{a}} \cos \theta \frac{\partial \mathrm{p}^{+}}{\partial \theta}=\mathrm{Q}_{2}  \tag{3.27}\\
\mathrm{~T}^{+}+\beta \Delta \mathrm{t}[\tau] \mathrm{D}^{+}=\mathrm{Q}_{3}  \tag{3.28}\\
\mathrm{q}^{+}=\mathrm{Q}_{4}  \tag{3.29}\\
\left(\ln \mathrm{P}_{\text {surf }}\right)^{+}+\beta \Delta \mathrm{t}[\mathrm{v}] \mathrm{D}^{+}=\mathrm{Q}_{5} \tag{3.30}
\end{gather*}
$$

where the right-hand sides $Q_{1}-Q_{5}$ include all the terms which have been computed on the grid, and $T_{v}{ }^{+}$replaces $\mathrm{T}^{+}$if $\mathrm{T}_{\mathrm{v}}$ is the spectral variable. Equations (3.26)-(3.30) have exactly the same form as Eqs. (2.40)-(2.44) of the Eulerian model and are solved in exactly the same way, by first transforming to spectral space. After finding the new spectral coefficients at time $(t+\Delta t)$, horizontal diffusion is also applied in the same way as for the Eulerian version.

The implementation of the time-filtering for the semi-Lagrangian model is identical to that for the Eulerian version, as described in Subsection 2.2.6

### 3.7 OPTIMIZATION OF VERTICALLY NON-INTERPOLATING SCHEME

In the 'vertically non-interpolating' scheme, the departure point of each modified trajectory lies on a model level. For the set of arrival points on each model level, it is of interest to determine the frequency distribution of the corresponding departure points. The results of an experiment run to collect these statistics led to a significant optimization of the code for the vertically non-interpolating scheme.

The statistics were obtained from a 10-day forecast using the model in its operational configuration: T213, 31 levels, with a 15 -minute timestep. The results are summarized in Table 3.1, which shows that the vast majority ( $99.67 \%$ overall) of modified trajectories are horizontal; no departure point was ever more than three model levels away from its corresponding arrival point.

The implication of these results is that a great deal of redundant calculation was being performed in the vertically non-interpolating scheme. For each horizontal modified trajectory, the interpolation of the horizontal winds in the trajectory calculation itself becomes two-dimensional rather than three-dimensional, as do the interpolations of
'right-hand side' terms at the mid point of the trajectory, while the additional interpolations to calculate terms of the form $\dot{\eta}^{*}(\partial X / \partial \eta)^{0}$ in (3.12) are not required at all. Consequently, special routines were written to perform interpolations which are two-dimensional everywhere except at a set of 'flagged' points where they become threedimensional, and similarly to perform two- or three-dimensional interpolations at the flagged points while skipping all other points. The use of these special routines reduced the 'semi-Lagrangian overhead' for the vertically noninterpolating scheme by about $30 \%$.

TABLE 3.1 FREQUENCY DISTRIBUTION (\%) OF DEPARTURE POINTS IN THE ‘VERTICALLY NON-INTERPOLATING’ SCHEME

| Arrival level K | Departure levels |  |  |  |
| :---: | :---: | :---: | :---: | :---: |
|  | k | $k \pm 1$ | $k \pm 2$ | $k \pm 3$ |
| 1-6 | 100.00 |  |  |  |
| 7-9 | 100.00 | * |  |  |
| 10 | 99.99 | 0.01 | * |  |
| 11 | 99.96 | 0.04 | * |  |
| 12 | 99.89 | 0.11 | * |  |
| 13 | 99.76 | 0.24 | * | * |
| 14 | 99.60 | 0.40 | * | * |
| 15 | 99.43 | 0.57 | * | * |
| 16 | 99.28 | 0.72 | * |  |
| 17 | 99.16 | 0.83 | 0.01 |  |
| 18 | 99.08 | 0.92 | * |  |
| 19 | 99.05 | 0.94 | 0.01 |  |
| 20 | 99.05 | 0.94 | 0.01 | * |
| 21 | 99.09 | 0.91 | * | * |
| 22 | 99.14 | 0.85 | 0.01 | * |
| 23 | 99.22 | 0.78 | * |  |
| 24 | 99.31 | 0.69 | * |  |
| 25 | 99.44 | 0.56 | * |  |
| 26 | 99.60 | 0.40 | * |  |
| 27 | 99.78 | 0.22 | * |  |
| 28 | 99.92 | 0.08 | * |  |
| 29 | 99.99 | 0.01 |  |  |
| 30 | 100.00 | * |  |  |
| 31 | 100.00 |  |  |  |

Asterisks indicate less than $0.005 \%$ frequency

### 3.8 MODIFIED SEMI-LAGRANGIAN EQUATIONS

### 3.8.1 Momentum equations

The momentum equations are treated in vector form (Eq. (3.14)). Following Rochas (1990) and Temperton (1997), the Coriolis terms can be incorporated in the semi-Lagrangian advection. Thus, the advected variable becomes ${\underset{\sim}{V}}_{H}+2 \underset{\sim}{\Omega} \times \underset{\sim}{r}$ where $\underset{\sim}{\Omega}$ is the earth's rotation and $\underset{\sim}{r}$ is the radial position vector, while the Coriolis terms are dropped from the right-hand side. As described by Temperton (1997), this reformulation is beneficial provided that the spherical geometry is treated accurately in determining the departure point and in rotating the vectors to account for the change in the orientation of the coordinate system as the particle follows the trajectory.

The discretization of the momentum equations in the notation of Eq. (3.1) is then:

$$
\begin{gather*}
X={\underset{\sim}{V}}_{H}+2 \underset{\sim}{\Omega} \times \underset{\sim}{r}  \tag{3.31}\\
L=-\underset{\sim}{\nabla}\left(\underset{\gamma}{\gamma} T_{v}+R_{\text {dry }} T_{\text {ref }} \ln p_{\text {surf }}\right)  \tag{3.32}\\
N=-\left(\underset{\sim}{\nabla} \phi+R_{\text {dry }} T_{v} \underset{\sim}{\nabla} \ln p\right)-L \tag{3.33}
\end{gather*}
$$

where $R_{d r y}$ is the gas constant for dry air, $T_{\text {ref }}$ is a reference temperature, $\phi$ is geopotential and $\underline{\gamma}$ is the linearized hydrostatic integration matrix defined in Eq. (2.32) of Ritchie (1995).

In component form, $2 \underset{\sim}{\Omega} \times \underset{\sim}{r}$ is just $(2 \Omega \operatorname{acos} \theta, 0)$ where $a$ is the earth's radius and $\theta$ is latitude. Since the latitude of the departure point is known, the term $2 \Omega \times \underset{\sim}{r}$ in the advected variable $X$ is computed analytically rather than interpolated. An alternative semi-implicit treatment of the Coriolis terms has also been developed (Temperton 1997).

### 3.8.2 Continuity equation

Modelling flow over mountains with a semi-Lagrangian integration scheme can lead to problems in the form of a spurious resonant response to steady orographic forcing. The mechanism was clarified by Rivest et al. (1994). Strictly speaking, the problem has little to do with the semi-Lagrangian scheme itself; rather, it is a result of the long time steps permitted by the scheme, such that the Courant number becomes greater than 1 . Recently, Ritchie and Tanguay (1996) proposed a modification to the semi-Lagrangian scheme which alleviates the problem. It turned out that their suggestion was easy to implement in the ECMWF model, and had additional benefits besides improving the forecast of flow over orography.

Although Ritchie and Tanguay start by introducing a change of variables in the semi-implicit time discretization, this is not necessary and a slightly different derivation is presented here. The continuity equation is written in the form

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{dt}}\left(\ln \mathrm{p}_{\text {surf }}\right)=[\mathrm{RHS}] \tag{3.34}
\end{equation*}
$$

where [RHS] represents right-hand-side terms. The total derivative on the left-hand side is discretized in a semiLagrangian fashion, and the final form of the discretized equation involves a vertical summation.

Now split $\ln \mathrm{p}_{\mathrm{s}}$ into two parts:

$$
\begin{equation*}
\ln \mathrm{p}_{\text {surf }}=\mathrm{I}^{*}+\mathrm{I}^{\prime} \tag{3.35}
\end{equation*}
$$

where the time-independent part I* depends on the underlying orography $\phi_{\text {surf }}$ :

$$
\begin{equation*}
I^{*}=\left(-\phi_{\text {surf }}\right) /\left(\mathrm{R}_{\text {dry }} \overline{\mathrm{T}}\right) \tag{3.36}
\end{equation*}
$$

and $\overline{\mathrm{T}}$ is a reference temperature. This choice gives

$$
\begin{equation*}
\nabla \phi_{\text {surf }}+\mathrm{R}_{\mathrm{dry}} \bar{\top} \nabla l^{*}=0 \tag{3.37}
\end{equation*}
$$

so that I* is (to within an additive constant) the value of $\ln \mathrm{p}_{\text {surf }}$ appropriate for an isothermal state at rest with underlying orography.

Using (3.35) and (3.36),

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{dt}}\left(\left(\ln \mathrm{p}_{\text {surf }}\right)\right)=\frac{\mathrm{d} \mathrm{l}^{\prime}}{\mathrm{dt}}-\left(\frac{1}{\left.\left.\mathrm{R}_{\mathrm{dry}} \overline{\bar{T}}{\underset{\mathrm{v}}{\mathrm{H}}} \cdot \underset{\sim}{\nabla} \phi_{\text {surf }}\right), ~\right) ~}\right. \tag{3.38}
\end{equation*}
$$

The second term on the right-hand side is computed in an Eulerian manner and transferred to the right-hand side of the continuity equation (3.34), which becomes

$$
\begin{equation*}
\frac{\mathrm{dl}^{\prime}}{\mathrm{dt}}=[R H S]+\frac{1}{\mathrm{R}_{\mathrm{dry}} \overline{\bar{T}}} \underline{\mathrm{v}}_{\mathrm{H}} \cdot \underset{\sim}{\nabla} \phi_{\text {surf }} \tag{3.39}
\end{equation*}
$$

The new advected variable is much smoother than the original, since the influence of the underlying orography has been subtracted out; hence, the semi-Lagrangian advection is presumably more accurate.

### 3.8.3 Thermodynamic equation

As mentioned above, the semi-Lagrangian treatment of the continuity equation is improved by changing the advected variable to a smoother quantity which is essentially independent of the underlying orography. A similar modification has been implemented in the thermodynamic equation, borrowing an idea from the treatment of horizontal diffusion. To approximate horizontal diffusion on pressure surfaces, thereby avoiding spurious warming over mountain tops in sigma or hybrid vertical coordinates, the diffused quantity is ( $T-T_{c}$ ), with

$$
\begin{equation*}
T_{c}=\left(p_{\text {surf }} \frac{\partial p}{\partial p_{\text {surf }}} \frac{\partial T}{\partial p}\right)_{\text {ref }} \ln p_{\text {surf }} \tag{3.40}
\end{equation*}
$$

where the subscript 'ref' denotes a reference value which is a function only of model level. For the purposes of the semi-Lagrangian advection $\ln \mathrm{p}_{\text {surf }}$ is replaced by a time-independent value as in Eq. (3.36) above, to define a "temperature" $\mathrm{T}_{\mathrm{b}}$ which depends only on the model level and the underlying orography:

$$
\begin{equation*}
T_{b}=-\left(p_{\text {surf }} \frac{\partial p}{\partial p_{\text {surf }}} \frac{\partial T}{\partial p}\right)_{\text {ref }} \cdot \phi_{\text {surf }} /\left(R_{\text {dry }} \bar{T}\right) \tag{3.41}
\end{equation*}
$$

The semi-Lagrangian advection is now applied to the quantity $\left(T-T_{b}\right)$, while a compensating expression

$$
\begin{equation*}
-{\underset{\sim}{\mathrm{v}}}_{\mathrm{H}} \cdot \nabla \mathrm{~T}_{\mathrm{b}}-\dot{\eta} \frac{\partial \mathrm{T}_{\mathrm{b}}}{\partial \eta} \tag{3.42}
\end{equation*}
$$

appears on the right-hand side of the equation and is computed in an Eulerian fashion (note that this time it includes a vertical advection term).

### 3.9 Two-TIME-LEVEL SEMI-LAGRANGIAN SCHEME

Formally, a two-time-level scheme may be written in the notation of Eq.(3.2) as:

$$
\begin{equation*}
\frac{\mathrm{X}_{\mathrm{A}}^{+}-\mathrm{X}_{\mathrm{D}}^{-}}{\Delta \mathrm{t}}=\frac{1}{2}\left(\mathrm{~L}_{\mathrm{D}}^{-}+\mathrm{L}_{\mathrm{A}}^{+}\right)+\frac{1}{2}\left(\mathrm{~N}_{\mathrm{D}}^{*}+\mathrm{N}_{\mathrm{A}}^{*}\right) \tag{3.43}
\end{equation*}
$$

where
$\mathrm{X}_{\mathrm{A}}^{+}=\mathrm{X}(\underset{\sim}{\mathrm{x}}, \mathrm{t}+\Delta \mathrm{t})$ is the value at the "arrival" gridpoint at $(\mathrm{t}+\Delta \mathrm{t})$
$\mathrm{X}_{\mathrm{D}}^{-}=\mathrm{X}(\underset{\sim}{\mathrm{x}}-\underset{\sim}{\alpha}, \mathrm{t}) \mathrm{s}$ the value interpolated at the "departure" point at time t ;
$\mathrm{L}_{\mathrm{A}}^{+}$and $\mathrm{L}_{\mathrm{D}}^{-}$are the linear terms defined similarly;
$N^{*}$ are the non linear terms, obtained by extrapolation in time to $\left(t+\frac{1}{2} \Delta t\right)$

$$
\begin{equation*}
\mathrm{N}^{*}=\frac{3}{2} \mathrm{~N}(\mathrm{t})-\frac{1}{2} \mathrm{~N}(\mathrm{t}-\Delta \mathrm{t}) \tag{3.44}
\end{equation*}
$$

The displacement equation becomes

$$
\begin{equation*}
\underset{\sim}{\alpha}=\Delta \mathrm{t}{\underset{\sim}{*}}^{*}\left(\underset{\sim}{x}-\frac{1}{2} \underset{\sim}{\alpha}, \mathrm{t}+\frac{1}{2} \Delta \mathrm{t}\right) \tag{3.45}
\end{equation*}
$$

where the three-dimensional wind field $\underset{\sim}{V}$ is also extrapolated in time:

$$
\begin{equation*}
{\underset{\sim}{V}}^{*}=\frac{3}{2} \underset{\sim}{V}(\mathrm{t})-\frac{1}{2} \underset{\sim}{\mathrm{~V}}(\mathrm{t}-\Delta \mathrm{t}) \tag{3.46}
\end{equation*}
$$

The iterative scheme and first-guess for solving (3.45) are exactly analogous to those for solving (3.5).
The choices for the variables X and for the interpolation schemes remain exactly as for the three-time-level scheme.

The semi-implicit equations to be solved in spectral space have the same form as for the three-time-level scheme, except that $\Delta \mathrm{t}$ is replaced by $\Delta \mathrm{t} / 2$.

In principle a two-time-level scheme should have no $2 \Delta t$ computational mode, and a time-filtering procedure is no longer needed.

### 3.9.1 Stable Extrapolation Two-Time-Level Scheme (SETTLS)

An alternative second-order accurate scheme to solving (3.45) can be derived by expanding the position vector $\vec{R}$ of the parcel of air as a Taylor series around the departure point of the semi-Lagrangian trajectory:

$$
\begin{equation*}
\overrightarrow{\mathrm{R}}_{\mathrm{A}}^{\mathrm{t}+\Delta \mathrm{t}} \approx \overrightarrow{\mathrm{R}}_{\mathrm{D}}^{\mathrm{t}}+\Delta \mathrm{t} \cdot\left[\frac{\mathrm{~d}}{\mathrm{dt}} \overrightarrow{\mathrm{R}}\right]_{\mathrm{D}}^{\mathrm{t}}+\frac{\Delta \mathrm{t}^{2}}{2} \cdot\left[\frac{\mathrm{~d}^{2}}{\mathrm{dt}^{2}}(\overrightarrow{\mathrm{R}})\right]_{\mathrm{AV}} \tag{3.47}
\end{equation*}
$$

Here subscript AV indicates some average value along the semi-Lagrangian trajectory.
Substituting the time derivative of $\vec{R}$ by the velocity vector $\vec{V}$, we find

$$
\begin{equation*}
\overrightarrow{\mathrm{R}}_{\mathrm{A}}^{\mathrm{t}+\Delta \mathrm{t}} \approx \overrightarrow{\mathrm{R}}_{\mathrm{D}}^{\mathrm{t}}+\Delta \mathrm{t} \cdot \overrightarrow{\mathrm{~V}}_{\mathrm{D}}^{\mathrm{t}}+\frac{\Delta \mathrm{t}^{2}}{2} \cdot\left[\frac{\mathrm{~d}}{\mathrm{dt}}(\overrightarrow{\mathrm{~V}})\right]_{\mathrm{AV}} \tag{3.48}
\end{equation*}
$$

This equation describes an uniformly accelerated movement. The trajectories can no longer be considered as straight lines on a plane or as arcs of a great circle in spherical geometry as is traditionally done in semi-Lagrangian schemes and the position of the middle point of the trajectory is no longer an average between the departure and the arrival points.

To proceed, one has to estimate the quantity

$$
\begin{equation*}
\left[\frac{\mathrm{d}}{\mathrm{dt}}(\overrightarrow{\mathrm{~V}})\right]_{\mathrm{AV}} \tag{3.49}
\end{equation*}
$$

To estimate (3.49) the first possibility explored was to use an average along the trajectory of the explicit estimate of the r.h.s. of the momentum equations as the horizontal part of expression (3.49) and the expression

$$
\begin{equation*}
\left[\frac{\mathrm{d}}{\mathrm{dt}}(\mathrm{~W})\right]_{\mathrm{AV}} \approx \frac{\mathrm{~W}_{\mathrm{A}}^{\mathrm{t}}-\mathrm{W}_{\mathrm{D}}^{\mathrm{t}-\Delta \mathrm{t}}}{\Delta \mathrm{t}} \tag{3.50}
\end{equation*}
$$

for the vertical part.
After exploring many other possibilities, the following estimate was adopted:

$$
\begin{equation*}
\left[\frac{\mathrm{d}}{\mathrm{dt}}(\overrightarrow{\mathrm{~V}})\right]_{\mathrm{AV}} \approx\left[\frac{\mathrm{~d}}{\mathrm{dt}}(\overrightarrow{\mathrm{~V}})\right]^{\mathrm{t}-\frac{\Delta \mathrm{t}}{2}} \approx \frac{\overrightarrow{\mathrm{~V}}_{\mathrm{A}}^{\mathrm{t}}-\overrightarrow{\mathrm{V}}_{\mathrm{D}}^{\mathrm{t}-\Delta \mathrm{t}}}{\Delta \mathrm{t}} \tag{3.51}
\end{equation*}
$$

using the departure point of the semi-Lagrangian trajectory corresponding to the present time step instead of the departure point of the trajectory corresponding to the previous time step. Here D means the position at time $t$ of the parcel of air which will arrive to gridpoint $A$ at time $t+\Delta t$.

This estimate assumes that the total time derivative of the velocity is constant with time, following Durran's suggestion of "extrapolating along the trajectory", but the estimate uses only the arrival and departure points of the present trajectory and is therefore compatible with the semi-implicit treatment of the evolution equations. This scheme should therefore be also stable according to linear stability analysis and has accordingly been named "Stable Extrapolation Two-Time-Level Scheme" or SETTLS.

Substituting (3.51) into (3.48) we obtain:

$$
\begin{equation*}
\overrightarrow{\mathrm{R}}_{\mathrm{A}}^{\mathrm{t}+\Delta \mathrm{t}}=\overrightarrow{\mathrm{R}}_{\mathrm{D}}^{\mathrm{t}}+\frac{\Delta \mathrm{t}}{2} \cdot\left(\left[2 \overrightarrow{\mathrm{~V}}^{\mathrm{t}}-\overrightarrow{\mathrm{V}}^{\mathrm{t}-\Delta \mathrm{t}}\right]_{\mathrm{D}}+\overrightarrow{\mathrm{V}}_{\mathrm{A}}^{\mathrm{t}}\right) \tag{3.52}
\end{equation*}
$$

and a similar expression can be used in every evolution equation to treat the non-linear terms of the r.h.s.

### 3.10 NUMERICAL COUPLING OF THE PHYSICAL PARAMETERIZATIONS TO THE 'DYNAMICAL" EQUATIONS (SLAVEPP)

Due to the diffusive nature of the mostly parabolic equations in the physics the contributions of the physical parameterizations are computed separately from the "dynamical" equations. The coupling of these two parts can use
 cussed in detail by Wedi (1999).

In equation (3.14)-(3.16) the contribution of the physical parameterizations are denoted as P indicating an evaluation of the parameterizations at the arrival point only. In the two time level scheme as described in section 3.9 this is replaced by a partly second order accurate coupling of the parameterizations in time and space, which is achieved by evaluating part of the "physics" at the arrival point and the remainder at the departure point of the semiLagrangian trajectory. Due to the different nature of the parameterized processes the contributions of radiation, convection and cloud parameterization are averaged "along" the semi-Lagrangian trajectory while the contributions of vertical diffusion and parameterized gravity waves are taken at the arrival point only. Equation (3.43) becomes then

$$
\begin{equation*}
\frac{\mathrm{X}_{\mathrm{A}}^{+}-\mathrm{X}_{\mathrm{D}}^{-}}{\Delta \mathrm{t}}=\frac{1}{2}\left(\mathrm{~L}_{\mathrm{D}}^{-}+\mathrm{L}_{\mathrm{A}}^{+}\right)+\frac{1}{2}\left(\mathrm{~N}_{\mathrm{D}}^{*}+\mathrm{N}_{\mathrm{A}}^{*}\right)+\frac{1}{2}\left(\mathrm{P}_{\mathrm{D}, \mathrm{rad}+\operatorname{conv}+\mathrm{cloud}}^{-}+\mathrm{P}_{\mathrm{A}, \mathrm{rad}+\mathrm{conv}+\operatorname{cloud}}^{+}\right)+\mathrm{P}_{\mathrm{A}, \mathrm{vdif}+\mathrm{gwdrag}}^{+} \tag{3.53}
\end{equation*}
$$

Part of the implicit calculations of the physical parameterizations use the following tendency:

$$
\begin{equation*}
\tilde{\mathrm{D}_{\mathrm{A}}}=\frac{\tilde{\mathrm{X}}_{\mathrm{A}, \mathrm{Dyn}}^{+}-\mathrm{X}_{\mathrm{A}}^{-}}{\Delta \mathrm{t}} \tag{3.54}
\end{equation*}
$$

with equation (3.43) modified to yield

$$
\begin{equation*}
\frac{\tilde{X}_{\mathrm{A}, \mathrm{Dyn}}^{+}-\mathrm{X}_{\mathrm{D}}^{-}}{\Delta \mathrm{t}}=\frac{1}{2}\left(\mathrm{~L}_{\mathrm{D}}^{-}+\tilde{\mathrm{L}}_{\mathrm{A}}^{+}\right)+\frac{1}{2}\left(\mathrm{~N}_{\mathrm{D}}^{*}+\mathrm{N}_{\mathrm{A}}^{*}\right) \tag{3.55}
\end{equation*}
$$

The " $\sim$ " denotes that only provisional values of the dynamic fields are available because semi-implicit correction terms are still to be computed (see section 3.6). Therefore $L_{A}^{+} \approx \tilde{L}_{A}^{+}=L_{A}^{-}$is used for the linear terms. Equation (3.54) describes local tendencies, which are computed subtracting the new provisional explicit values $\tilde{X}_{A}^{+}$, Dyn of the dynamic fields (at the arrival point) from their values $X_{A}^{-}$at the previous time step. The parameterizations at the time step $t+\Delta t$ are computed at the arrival point as shown in the following equation:

$$
\begin{align*}
& \mathrm{P}_{\mathrm{A}}^{+}=\mathrm{P}_{\mathrm{A}, \text { rad }}^{+}\left(\mathrm{X}_{\mathrm{A}}^{-}\right)  \tag{3.56}\\
& +P_{A, \text { vdif }}^{+}\left(X_{A}^{-}, \tilde{D_{A}}, P_{A, \text { rad }}^{+}\right) \\
& +\mathrm{P}_{\mathrm{A}, \mathrm{gwdrag}}^{+}\left(\mathrm{X}_{\mathrm{A}}^{-}, \mathrm{D}_{\mathrm{A}}\right) \\
& +\mathrm{P}_{\mathrm{A}, \text { conv }}^{+}\left(\mathrm{X}_{\mathrm{A}, \text { predict }}^{+} \text {(output) } \mathrm{F}_{\text {conv }}\right) \\
& +\mathrm{P}_{\mathrm{A}, \text { cloud }}^{+}\left(\mathrm{X}_{\mathrm{A}, \text { predict }}^{+}(\text {input }) \mathrm{F}_{\text {conv }}\right)
\end{align*}
$$

where the "first guess" predictor $\mathrm{X}_{\mathrm{A}, \text { predict }}$ of the model variables at the arrival point at time step $\mathrm{t}+\Delta \mathrm{t}$ is computed from the tendency of the "dynamics", the tendency of the parameterizations of radiation, convection and clouds at the previous time-step $t$ and the tendency of vertical diffusion and gravity waves at $\mathrm{t}+\Delta \mathrm{t}$ :

$$
\begin{equation*}
\mathrm{X}_{\mathrm{A}, \text { predict }}^{+}=\tilde{\mathrm{X}}_{\mathrm{A}, \mathrm{Dyn}}^{+}+\alpha \mathrm{P}_{\mathrm{D}, \text { rad }+ \text { conv }+ \text { cloud } \Delta \mathrm{t}+\mathrm{P}_{\mathrm{A}, \text { vdif }+ \text { gwdrag }}^{-} \Delta \mathrm{t} . ~}^{\text {. }} \tag{3.57}
\end{equation*}
$$

$F_{\text {conv }}$ denotes an explicit interaction of the parameterizations of cloud and convection. The parameter $\alpha=0.5$ has been introduced in order to achieve a better balance between the physical parameterizations when the "first guess" predictor is computed.

### 3.10.1 Moisture adjustment and first time-step treatment

The parametrizations of cloud and convection show a sensitivity to the initial profiles. Therefore, at the initial timestep of the model "first guess" predictors are generated by a two step iteration of the parametrizations of cloud and convection consistent with the provisional dynamic fields as described above.

The effective profiles of temperature and humidity (including all contributions from the departure as well as the arrival point) are computed after all physical processes have been accounted for. A final moist adjustment is performed on these effective profiles and any amount of surplus humidity is added to the rainfall or snow fluxes in the next time-step. Note, that after this adjustment the temperature profile may still be altered as a result of the semiimplicit solution procedure.

Part III: Dynamics and numerical procedures

## CHAPTER 4 Computational details

Implementing a high-resolution model, which must run operationally within a given elapsed time on a given computer system, presents a number of interesting technical challenges. In this section we present some of the computational details which enabled the goal of implementation to be achieved.

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4.1 Scanning structure
4.2 Multitasking
4.3 Performance
4.4 Analysis of CPU time
4.5 The IFS model

### 4.1 SCANNING STRUCTURE

Each timestep of the model integration procedure consists of three scans.
At the beginning of the timestep, the model fields at time $t-\Delta t$ are specified in grid-point form (as described in Subsection 2.2.7 of Chapter 2 'Basic equations and discretization', these fields are 'partially time-filtered'). The grid-point values of the model variables are contained in a 'grid-point work-file', held on a secondary storage device and organized as a random-access file with one record for each latitude row. Meanwhile, the model fields at time t are specified in spectral form, all the spectral coefficients being held in central memory. The first scan consists of Legendre transforms to compute the Fourier coefficients of the model variables at time $t$ on each latitude row, the results being written out to a 'Fourier work-file', again organized with one record for each latitude row. During the first scan, latitude rows are processed in north/south pairs with the members of each pair being equidistant from the equator, in order to make use of the symmetries of the Legendre polynomials (see for example Temperton,1991). Once this first scan has been completed, the spectral coefficients are no longer required and the central memory arrays can be released for use during the next scan.

The second scan steps through the latitude rows, starting at the row nearest the North Pole and proceeding southwards. At each row, the corresponding records of the grid-point values at time ( $\mathrm{t}-\Delta \mathrm{t}$ ) and the Fourier coefficients at time $t$ are read in. Fourier transforms then provide grid-point values of the fields (together with any required horizontal derivatives) at time t . At this juncture, the time-filtering of the fields at time $(\mathrm{t}-\Delta \mathrm{t})$ is completed, while 'partially time-filtered' fields at time $t$ are also computed and written out to the grid-point work-file ready for the next timestep.

The grid-point calculations for the present timestep continue using the time-filtered values at $(t-\Delta t)$ and the unfiltered values at time $t$. The right-hand sides of the equations, discretized in semi-Lagrangian form as described in Chapter 3, are computed with terms being grouped separately depending on whether they will be evaluated at the departure point, the midpoint or the arrival point of the trajectory. The results of these calculations, together with the horizontal wind components and the vertical velocity $\dot{\eta}$, are then stored in a 'rotating buffer' which contains values for a number of consecutive latitude rows. The grid-point calculations described so far correspond to
the southernmost row contained in this buffer. Next, the focus of the computation returns to the central row of the buffer. Values of the wind fields and the right-hand sides of the equations are now available at a sufficient distance to the north and south of the central row for the trajectory calculations to be performed and for the semi-Lagrangian timestep to be implemented, thus furnishing provisional values at $(\mathrm{t}-\Delta \mathrm{t})$. As described in Section 3.4, the contributions from the physical parametrization schemes can then be incorporated to complete the calculation of the right-hand sides $\mathrm{Q}_{1}-\mathrm{Q}_{5}$ of Eqs. (3.26)-(3.30).

These right-hand sides are now Fourier transformed and the coefficients are written out to another Fourier workfile, again organized with one record for each latitude row but this time with a special structure which will be exploited in the third scan. The computation then proceeds southwards to the next pair of 'southernmost' and 'central' rows, the values computed for the new southern row overwriting those in the buffer for the previous northernmost row, which are no longer required.

At the start of the second scan, there is clearly an initialization phase during the first few rows when only the first part of the above calculations can be done. Similarly, at the end of the scan there is a 'winding-down' phase during which the first part of the calculations has already been done, and only the second part is required. The same logical structure is also used to run the Eulerian version of the model, but in this case the width of the 'rotating buffer' can be reduced to that for a single latitude row.

The third scan performs direct Legendre transforms to obtain the provisional spectral coefficients at time ( $\mathrm{t}+\Delta \mathrm{t}$ ) from the Fourier coefficients computed in the second scan, using Gaussian quadrature. The calculation proceeds one zonal wavenumber at a time. Here we make use of the special structure of the Fourier work-file; although the file was written row by row, it can be read in 'transposed' fashion, wavenumber by wavenumber. The direct Legendre transforms first exploit the symmetries of the Legendre polynomials, and then complete the calculations using highly efficient matrix multiplication routines. To see how this is achieved, notice that since a single Legendre transform can be written as a matrix/vector multiplication of the form $\underset{\sim}{y}=[P] \underset{\sim}{x}$, a set of simultaneous transforms for the same zonal wavenumber but for different variables and model levels can be written as

$$
\left\{{\underset{\sim}{x}}_{1}{\underset{\sim}{x}}_{2} \ldots{\underset{\sim}{y}}_{\mathrm{y}}\right\}=[\mathrm{P}]\left\{{\underset{\sim}{x}}_{1}{\underset{\sim}{2}}_{2} \ldots{\underset{\sim}{x}}_{\mathrm{x}}\right\}
$$

which is indeed in the form of a matrix multiplication $[Y]=[P][X]$. A similar technique could have been used in the first scan, and this has been incorporated in the latest version of the model.

After the transformation to spectral space, the semi-implicit equations are solved and the horizontal diffusion is implemented as described in Section 3.4, thus completing the calculation of the spectral coefficients at time $(t+\Delta t)$. At the end of the third scan, the whole model has been advanced by one timestep.

### 4.2 Multitasking

Currently the model is run on a 'modestly parallel' supercomputer (specifically, a Cray Y-MP C90 with 16 processors), and multitasking is an important aspect of the strategy to make the best use of the available computer power. We have chosen to rely mainly on high-level 'macrotasking', i.e., dividing the computation into large independent units of work, each of which is assigned to one of the processors. Here only a brief outline will be given; additional details and discussion are provided by Dent (1992).

In the first scan, the unit of work is a pair of latitude rows. Each pair is independent of all the others, and a simple dynamic scheduling technique can be used: as each processor becomes free, a new pair of rows is assigned to it.

In the second scan, the unit of work is a single latitude row. For the semi-Lagrangian version, the calculations for each row are no longer independent of those for all the other rows. The trajectory calculations and semi-Lagrangian
advection algorithm for the central row of the rotating buffer can only be carried out once the required calculations have been completed for all the neighbouring rows, and somewhat complex logic is required to control the multitasking during this scan.

In the third scan, the unit of work is a single zonal wavenumber. Each wavenumber is independent of all the others, and the scheduling technique used in the first scan can again be used. The work content of each wavenumber varies from a maximum at $\mathrm{m}=0$ to a minimum at the largest value of m (the 'tip' of the triangular truncation), and the dynamic scheduling technique is effective in spreading the work over available processors.

### 4.3 Performance

The following performance figures relate specifically to the operational version of the model run at horizontal resolution T213 with 31 levels on the 16-processor Cray Y-MP C90. With the model time-step set at 15 minutes the total CPU time per forecast day would be about 1.5 hours on a single processor, the corresponding elapsed time (excluding the post-processing) being 7 minutes when the work is shared amongst 16 processors. This represents a sustained computation speed of about 3.5 gigaflops $3.5 \times 10^{9}$ floating-point operations per second). The memory requirements are 49 Mwords of central memory plus 70 Mwords of secondary storage. Multitasking using 16 processors provides a speed-up factor of 13 compared with using a single processor. A typical operational 10-day forecast, including all the post-processing, takes 2 hours of elapsed time.

### 4.4 Analysis of CPU time

In developing a high-resolution spectral model, the cost of the transforms (particularly the Legendre transforms) may be a cause for concern (e.g. Côté and Staniforth, 1990). In the case of a semi-Lagrangian model, it is clearly important that the gain obtained through the use of longer time steps is not outweighed by the extra cost of the semiLagrangian scheme. In view of these concerns, it is of interest to analyse the CPU time required for our model. Table 4.1 shows the percentage breakdown for the Eulerian version, for the fully interpolating semi-Lagrangian scheme and for the vertically non-interpolating scheme, at T213/L31 resolution.

Table 4.1 Analysis of CPU time (\%)

|  | Eulerian | Fully <br> interpolating <br> semi-Lagrangian | Vertically non- <br> interpolating <br> semi-Lagrangian |
| :--- | :---: | :---: | :---: |
| Dynamics | 21 | 15 | 17 |
| Physics | 53 | 42 | 45 |
| FFT | 6 | 3 | 4 |
| Legendre <br> transforms <br> Semi-Lagrangian | 20 | 13 | 14 |

This analysis suggests that the spectral method is still perfectly viable at this resolution, and that considerably higher resolutions can be achieved before the cost of the transforms becomes a matter for serious concern. The overhead of the semi-Lagrangian scheme, particularly the non-interpolating version, is also quite modest; for the present resolution it permits a timestep of 15 minutes compared with 3 minutes for the Eulerian version, and the resulting reduction in the CPU time for the forecast is about a factor of four. The semi-Lagrangian overhead is in fact slightly
less than suggested by the figures in Table, since there is a simultaneous reduction in the number of transforms compared with the Eulerian scheme. Comparing the two variants of the semi-Lagrangian scheme, the overall CPU time for the non-interpolating version is $8.5 \%$ less than that for the fully interpolating version.

### 4.5 The IFS Model

On 2nd March 1994, the model code described above was replaced in operations by the IFS ('Integrated Forecasting System) model, developed in collaboration with Météo-France (where it is known as ARPEGE; see Courtier et al. (1991) for an account of this project). The new code includes all the features required for three- and fourdimensional variational data assimilation (Thépaut and Courtier, 1991; Rabier and Courtier, 1992), and for determining optimal unstable perturbations for ensemble prediction (Buizza et al. 1993). The computational structure of the forecast model component of the system is similar to that described above but includes further improvements in efficiency, notably the matrix-multiplication treatment of the Legendre transforms in the first scan as well as the third scan (see Subsection 4.1 above), and the option to combine several latitude rows together (for example near the poles of the reduced grid) resulting in longer vectors.

## Part III: Dynamics and numerical procedures

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